

Gravitational redshift in quantum-clock interferometry

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Relativistic effects in macroscopically delocalized quantum superpositions

- Macroscopically delocalized quantum superpositions: coherent superposition of atomic wave packets
 Kovachy et al., Nature (2015)
- Same relativistic effects on superposition components
 STANFORD UNIVERSITY (e.g. atomic clocks)
- ★ <u>Goal</u> (QM + GR): experiment with general relativistic effects acting *non-trivially* on the quantum superposition

Proper time as which-way information

 Quantum superposition of clocks (COM + internal state) experiencing different proper times



Zych et al., Nat. Comm. (2011)

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Outline

- I. Relativistic effects in macroscopically delocalized quantum superpositions
- 2. Key elements of quantum-clock interferometry
- 3. Major challenges in quantum-clock interferometry
- 4. Doubly differential scheme for gravitational-redshift measurements
- 5. Feasibility and extensions

Key elements of quantum-clock interferometry

Quantum-clock model



• Initialization pulse:

$$|\mathbf{g}\rangle \rightarrow |\Phi(0)\rangle = \frac{1}{\sqrt{2}} (|\mathbf{g}\rangle + i e^{i\varphi} |\mathbf{e}\rangle)$$

• Evolution:

$$\left| \Phi(\tau) \right\rangle \propto \frac{1}{\sqrt{2}} \Big(|\mathbf{g}\rangle + i \, e^{i\varphi} e^{-i\Delta E \, \tau/\hbar} |\mathbf{e}\rangle \Big)$$

• Quantum overlap: $\left| \langle \Phi(\tau_b) | \Phi(\tau_a) \rangle \right| = \cos\left(\frac{\Delta E}{2\hbar} \left(\tau_b - \tau_a \right) \right)$

• Comparison of independent clocks (after read-out pulse):



$$\Delta \tau_b - \Delta \tau_a \approx \left(g L_z / c^2\right) \Delta t$$

for optical atomic clocks $\Delta E \sim 1 \, {
m eV} \qquad L_z \sim 1 \, {
m cm}$

Instead of independent clocks we pursue a quantum superposition at different heights.

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Instead of independent clocks we pursue a quantum superposition at different heights.

- Theoretical description of the clock
 - two-level atom (internal state):

$$\hat{H} = \hat{H}_1 \otimes |\mathbf{g}\rangle \langle \mathbf{g}| + \hat{H}_2 \otimes |\mathbf{e}\rangle \langle \mathbf{e}|$$

$$m_1 = m_g$$

 $m_2 = m_g + \Delta m$
 $\Delta m = \Delta E/c^2$

classical action for COM motion:

$$S_n[x^{\mu}(\lambda)] = -m_n c^2 \int d\tau = -m_n c \int d\lambda \sqrt{-g_{\mu\nu}} \frac{dx^{\mu}}{d\lambda} \frac{dx^{\nu}}{d\lambda} \qquad (n = 1, 2)$$
free fall
$$S_n[x^{\mu}(\lambda)] = -m_n c^2 \int d\tau - \int d\tau V_n(x^{\mu}) \qquad \text{including}$$
external forces

Atom interferometry in curved spacetime (including relativistic effects)

- Wave-packet evolution in terms of
 - central trajectory (satisfies classical e.o.m.) $X^{\mu}(\lambda)$
 - centered wave packet $|\psi_{\rm c}^{(n)}(\tau_{\rm c})\rangle$
- Fermi-Walker frame associated with the central trajectory
 - valid for freely falling wave packet (geodesic)
 - but also with external forces / guiding potential (accel. trajectory)
 - approximately non-relativistic dynamics for centered wave packet



• Metric in *Fermi-Walker* coordinates: $X^{\mu}(\tau_{c}) = (c \tau_{c}, \mathbf{0})$

$$ds^{2} = g_{\mu\nu}dx^{\mu}dx^{\nu} = g_{00}c^{2}d\tau_{c}^{2} + 2g_{0i}c\,d\tau_{c}\,dx^{i} + g_{ij}\,dx^{i}dx^{j}$$

$$g_{00} = -(1 + \delta_{ij} a^{i}(\tau_{c}) x^{j}/c^{2})^{2} - R_{0i0j}(\tau_{c}, \mathbf{0}) x^{i}x^{j} + O(|\mathbf{x}|^{3})$$

$$g_{0i} = -\frac{2}{3}R_{0jik}(\tau_{c}, \mathbf{0}) x^{j}x^{k} + O(|\mathbf{x}|^{3})$$

$$g_{ij} = \delta_{ij} - \frac{1}{3}R_{ikjl}(\tau_{c}, \mathbf{0}) x^{k}x^{l} + O(|\mathbf{x}|^{3})$$

• Expanding the action for the centered wave packet:

$$S_n[\mathbf{x}(t)] \approx \int d\tau_{\rm c} \left[-m_n c^2 - V_n(\tau_{\rm c}, \mathbf{0}) + \frac{m_n}{2} \mathbf{v}^2 - \frac{1}{2} \mathbf{x}^{\rm T} \left(\mathcal{V}^{(n)}(\tau_{\rm c}) - m_n \Gamma(\tau_{\rm c}) \right) \mathbf{x} - V_{\rm anh.}^{(n)}(\tau_{\rm c}, \mathbf{x}) \right]$$

• Hamiltonian: $\hat{H}_n = m_n c^2 + V_n(\tau_c, \mathbf{0}) + \hat{H}_c^{(n)}$

$$\hat{H}_{c}^{(n)} = \frac{1}{2m_{n}}\,\hat{\mathbf{p}}^{2} + \frac{1}{2}\,\hat{\mathbf{x}}^{T}\left(\mathcal{V}^{(n)}(\tau_{c}) - m_{n}\Gamma(\tau_{c})\right)\hat{\mathbf{x}} \qquad \qquad \mathcal{V}_{ij}^{(n)}(\tau_{c}) = \left.\partial_{i}\partial_{j}V_{n}(\tau_{c},\mathbf{x})\right|_{\mathbf{x}=\mathbf{0}}$$

- Wave-packet evolution: $|\psi^{(n)}(\tau_{\rm c})\rangle = e^{iS_n/\hbar} |\psi^{(n)}_{\rm c}(\tau_{\rm c})\rangle$
 - propagation phase

$$\mathcal{S}_n = -\int_{\tau_1}^{\tau_2} d\tau_{\rm c} \left(m_n c^2 + V_n(\tau_{\rm c}, \mathbf{0}) \right)$$

centered wave packet

$$i\hbar \frac{d}{d au_{
m c}} \left| \psi_{
m c}^{(n)}(au_{
m c})
ight
angle = \hat{H}_{
m c} \left| \psi_{
m c}^{(n)}(au_{
m c})
ight
angle$$

• Full interferometer (including laser kicks):



• Detection probability at the exit port(s): $\langle \psi_{I} | \psi_{I} \rangle$

$$\langle \psi_{\mathrm{I}} | \psi_{\mathrm{I}}
angle = rac{1}{2} ig(1 + \cos \delta \phi ig)$$

• Phase shift: $\delta \phi = \phi_b - \phi_a + \delta \phi_{sep}$

Major challenges in quantum-clock interferometry

Insensitivity to gravitational redshift (in a uniform field)

• Consider a freely falling frame:



• Proper-time difference between the two interferometer branches \longrightarrow independent of g

(small dependence due to pulse timing suppressed by $(v_{\rm rec}/c) \sim 10^{-10}$)

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Differential recoil

• Different recoil velocities \rightarrow different central trajectories



 Implied changes of proper-time difference are comparable to signal of interest.

Small visibility changes

 Reduced interference visibility due to deceasing quantum overlap of clock states:

$$\left\langle \Psi_{\mathrm{I}} | \Psi_{\mathrm{I}} \right\rangle = \frac{1}{2} + \frac{1}{2} \left| \left\langle \Phi(\tau_{b}) | \Phi(\tau_{a}) \right\rangle \right| \cos \delta \phi \qquad \left| \left\langle \Phi(\tau_{b}) | \Phi(\tau_{a}) \right\rangle \right| = \cos \left(\frac{\Delta E}{2\hbar} \left(\tau_{b} - \tau_{a} \right) \right)$$

• Small effect for feasible parameter range:

 $\left| \langle \Phi(\tau_b) | \Phi(\tau_a) \rangle \right| = \cos\left(\frac{\omega_0}{2} \frac{g \,\Delta z}{c^2} \,\Delta t\right) \approx 1 - \left(10^{-3}\right)^2 / 2 \qquad \qquad \Delta E / \hbar = \omega_0 \approx 2\pi \times 4 \times 10^2 \,\mathrm{THz}$ $\Delta z = 1 \,\mathrm{cm}$ $\Delta t = 1 \,\mathrm{s}$

• Extremely difficult to measure such small changes of visibility, which are masked by other effects leading also to loss of visibility.

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Doubly differential scheme for gravitational-redshift measurement

Differential phase-shift measurement

Detection probability at first exit port (independent of internal state):

- Phase-shift difference directly related to visibility reduction.
- Precise differential phase-shift measurement involving state-selective detection is much more viable.

(*immune* to spurious loss of contrast + common-mode rejection of phase noise)

Two-photon pulse for clock initialization

• Level structure for group-II-type atoms (e.g. Sr, Yb) employed in optical atomic clocks:



- Two-photon process resonantly connecting the two clock states.

Two-photon pulse for clock initialization

• Level structure for group-II-type atoms (e.g. Sr, Yb) employed in optical atomic clocks:



- Two-photon process resonantly connecting the two clock states.
- Equal-frequency counter-propagating laser beams in lab frame:
 constant effective phase -> simultaneity hypersurfaces in lab frame

 $e^{-i\omega t}e^{i\mathbf{k}\cdot\mathbf{x}} \times e^{-i\omega t}e^{-i\mathbf{k}\cdot\mathbf{x}} = e^{-i\,2\omega t}$

Laboratory frame

• Compare differential phase-shift measurements for different initialization times:



$$\left(\delta\phi^{(2)}(t_{i}') - \delta\phi^{(1)}(t_{i}')\right) - \left(\delta\phi^{(2)}(t_{i}) - \delta\phi^{(1)}(t_{i})\right) = \frac{\Delta E}{2\hbar} \left(\Delta\tau_{b} - \Delta\tau_{a}\right) = \Delta m g \,\Delta z \,(t_{i}' - t_{i})/\hbar$$

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Laboratory frame

• Compare differential phase-shift measurements for different initialization times:



$$\left(\delta\phi^{(2)}(t'_{i}) - \delta\phi^{(1)}(t'_{i})\right) - \left(\delta\phi^{(2)}(t_{i}) - \delta\phi^{(1)}(t_{i})\right) = \frac{\Delta E}{2\hbar} \left(\Delta\tau_{b} - \Delta\tau_{a}\right) = \Delta m g \,\Delta z \,(t'_{i} - t_{i})/\hbar$$

Freely falling frame

• Relativity of simultaneity: $\Delta \tau_{\rm c} \approx -v(t) \Delta z/c^2 = g(t - t_{\rm ap}) \Delta z/c^2$



 $\left(\delta\phi^{(2)}(t_{\rm i}') - \delta\phi^{(1)}(t_{\rm i}')\right) - \left(\delta\phi^{(2)}(t_{\rm i}) - \delta\phi^{(1)}(t_{\rm i})\right) = \frac{\Delta E}{2\hbar} \left(\Delta\tau_b - \Delta\tau_a\right) = \Delta m \, g \, \Delta z \, (t_{\rm i}' - t_{\rm i})/\hbar$

Challenges addressed

- Comparing measurements with different initialization times
 sensitive to gravitational redshift + further immunity
- Almost no recoil from *initialization pulse*, small residual recoil with no impact on gravitational redshift measurement,

effect of differential recoil from second pair of Bragg pulses cancels out in doubly differential measurement. • Residual recoil with no influence on the phase-shift for the excited state:



Feasibility and extensions

Feasible implementation



HITec (Hannover)

- 10-m atomic fountains operating with Sr, Yb in Stanford & Hannover respectively.
- More than $2 \, \mathrm{s}$ of free evolution time.
- Doubly differential phase shift of $1 \mod 6$

 $\Delta E/\hbar = \omega_0 \approx 2\pi \times 4 \times 10^2 \,\mathrm{THz}$

 $\Delta z = 1 \,\mathrm{cm}$

 $\Delta t_{\rm i} = 1\,{\rm s}$

- Resolvable in a single shot for atomic clouds with $N = 10^6$ atoms (shot-noise limited)
- More compact set-ups possible with guided or hybrid interferometry (less mature).

Conclusion

- Measurement of relativistic effects in macroscopically delocalized quantum superpositions with quantum-clock interferometry.
- Important *challenges* in quantum-clock interferometry and its application to gravitational-redshift measurement.
- Promising <u>doubly differential scheme</u> that overcomes them.
- Feasible implementation in facilities soon to become operational.
- Applicable also to more compact set-ups based on guided or hybrid interferometry.

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Proper time in atom interferometers: Diffractive vs. specular mirrors

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Diffraction of atoms in internal-state superpositions

BUT high laser power required due to large detuning.

- Alternative diffraction mechanism based on simultaneous pair of single-photon transitions.
 - Applicable to *fermionic* isotopes such as ${}^{87}Sr$ and ${}^{171}Yb$.
 - Required lasers already available in (some of) those facilities.

• Sequence of simultaneous pairs of pulses:



• Net result:

- internal state unchanged
- momentum transfer: twice single-photon momentum
- equal-amplitude superposition: undiffracted + diffracted wave packet

• Sequence of simultaneous pairs of pulses:



 $\pi/2$ pulse

 π pulse

Same ac Stark shifts for both internal states

Any light shifts cancel out in the doubly differential measurement (provided that the laser intensities are stable).

Extension to guided interferometry

In principle, guided interferometry can be sensitive to the gravitational redshift:



• Nevertheless, the *doubly differential* measurement scheme has many advantages.

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In principle, guided interferometry can be sensitive to the gravitational redshift:



• Nevertheless, the *doubly differential* measurement scheme has many advantages.

Extension to hybrid interferometers

 Intermediate stage with atoms held in an optical lattice, where they undergo Bloch oscillations:



• Similarly to pure *light-pulse* atom interferometers, they are insensitive to the gravitational redshift.

Extension to hybrid interferometers

• Intermediate stage with atoms held in an optical lattice, where they undergo Bloch oscillations:



• The *doubly differential* scheme can also be employed for measuring the gravitational redshift.

Other aspects

Proper-time difference in open interferometers



Proper-time difference in open interferometers



Proper-time difference and gravity gradients

