

First results from a test of Lorentz Invariance with the MICROSCOPE mission

- ACES Workshop 2018 -

Hélène Pihan-Le Bars¹, C. Guerlin^{2,1}, Quentin Bailey³, Geoffrey Mo⁴, Jay Tasson⁴, Peter Wolf¹

1 SYRTE, Observatoire de Paris, Université PSL, CNRS, Sorbonne Université, LNE, France

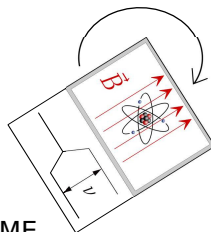
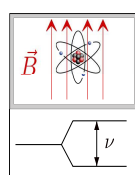
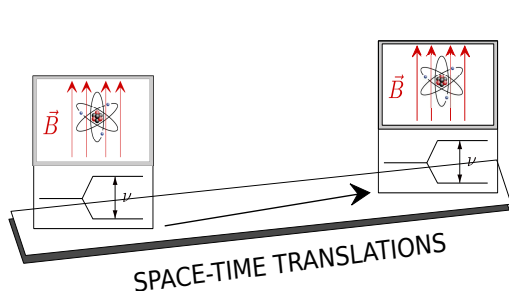
2 LKB, ENS, Université PSL, CNRS, Sorbonne Université, Collège de France, France

3 Embry-Riddle Aeronautical University, Arizona, USA

4 Carleton College, Minnesota, USA



Space-time symmetries

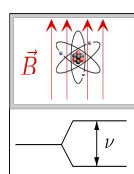
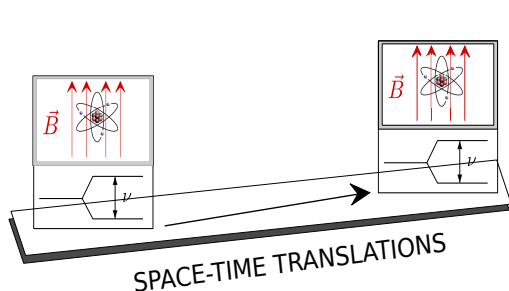


Homogeneous and isotropic space-time

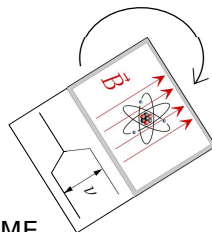
Physical laws are the same:

- ◇ whatever you choose as spacetime coordinates → **observer invariance**
- ◇ whatever boost or orientation changes you realize in a given inertial frame → **particle invariance**

Space-time symmetries



SPACE-TIME
ROTATIONS



Homogeneous and isotropic space-time

Physical laws are the same:

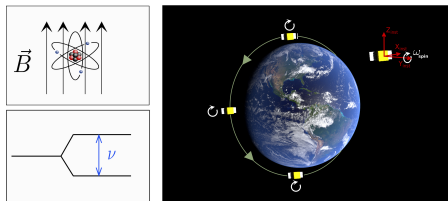
- whatever you choose as spacetime coordinates → **observer invariance**
- **whatever boost or orientation changes you realize in a given inertial frame → particle invariance**

The Lorentz symmetry group

Local Lorentz Invariance Principle

The results of a non-gravitational experiment is independent of the boost and orientation changes encountered by the free-falling laboratory frame.

C. Will, 1993



Lorentz symmetry in physics

- ◇ General Relativity → directly related to space-time geometry
- ◇ Standard Model → affect the space-time in which particles interact

Lorentz symmetry in alternative theories

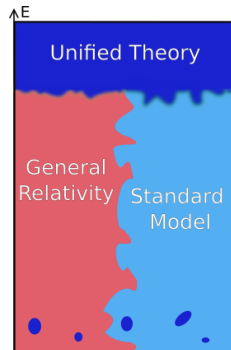
Some theories beyond General Relativity and Standard Model consider the possibility of Lorentz symmetry breaking at high energy (e.g. some string theories, Horava gravity...)

High energy effects (10^{19} GeV)

- Not accessible by experiments
- No low-energy experimental observable

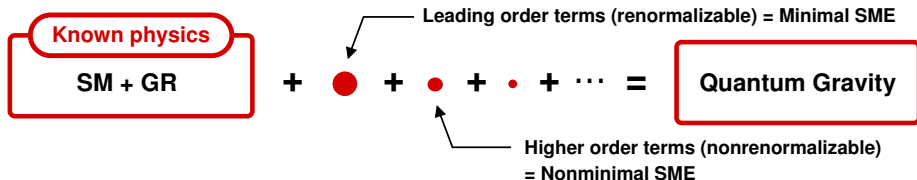
Solution: the Standard Model Extension (SME)

- Research of tiny low-energy signatures
- General theoretical framework: enable to derive observables for a large range of experiments

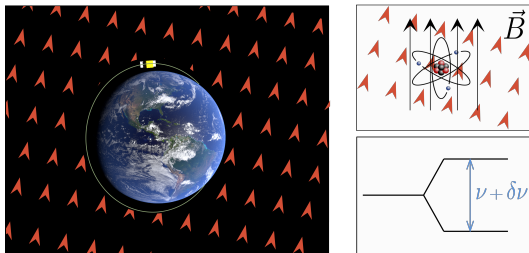


The Standard Model Extension

- Effective theory parametrizing all possible particle Lorentz violations
- Build from SM fields (fermions, photon, quark... sectors)
- Include violations arising from spacetime metric and curvature through a gravitational sector
- Amplitude of violations quantified by SME coefficients corresponding to components of SME tensors



The Standard Model Extension



- SME tensors are fixed in space and time and define a "preferred orientation" of space-time
- CONVENTION: SME coefficients expressed in Sun Centered Frame (inertiel over thousand of years)

Kostelecky et al., PRD 51, 1995, Kostelecky et al., PRD 58, 1998

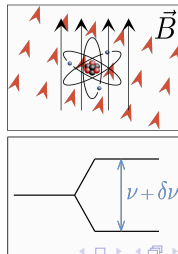
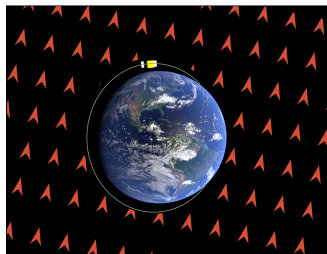
How to test Lorentz invariance

Requirements

- Experiment with a preferred orientation
- Experiment with long term stability

Signals: shape and dependance

- Search of periodic signals related to the experiment orientation changes with respect to SME background fields
 - ◇ spatial orientation
 - ◇ boost $\beta = v/c$

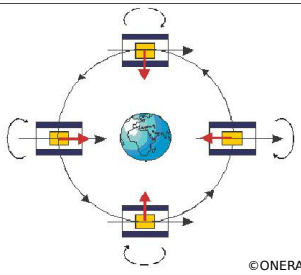
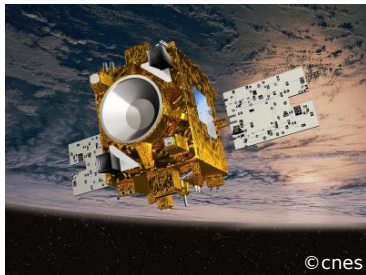


MICROSCOPE mission

Goal

Test of universality of free fall (UFF), with a precision of 10^{-15} on the Eötvös parameter δ .

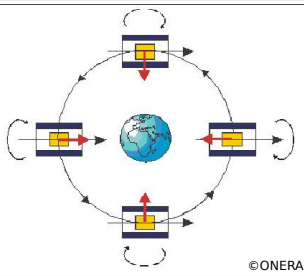
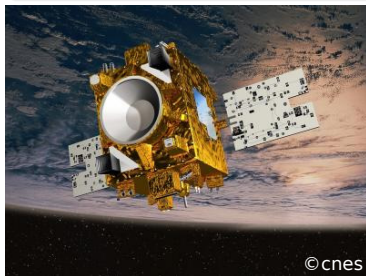
$$\delta(A, B) \sim \frac{m_{g,1}}{m_{i,1}} - \frac{m_{g,2}}{m_{i,2}}$$



MICROSCOPE mission

Instrument and measurements

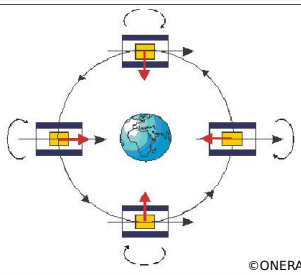
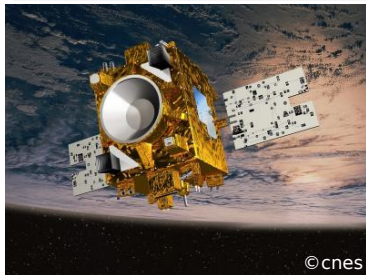
- Twin electrostatic accelerometer T-SAGE
- Measures the differential acceleration of two test masses (Ti/Pt) ou (Pt/Pt)



MICROSCOPE mission

Satellite

- Micro-satellite provided by CNES
- Sun-synchronous circular orbit (altitude ~ 700 km)
- Launched in Avril 2016, de-orbiting ongoing



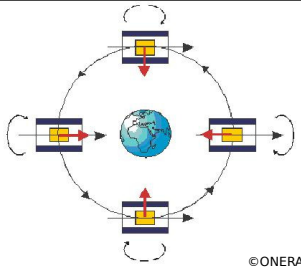
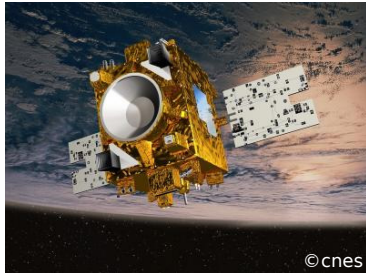
MICROSCOPE mission

First results

Obtained by the analysis of 120 orbits:

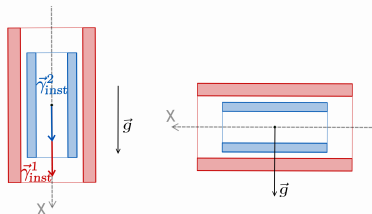
$$\delta = [-1 \pm 9 \text{ (stat)} \pm 9 \text{ (syst)}] \times 10^{-15}$$

Touboul et al., PRL, 119, 231101 (2017)



Relation between WEP and SME test

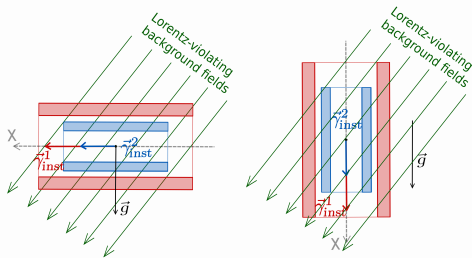
WEP violation



Signal $\neq 0$

No signal

SME violation

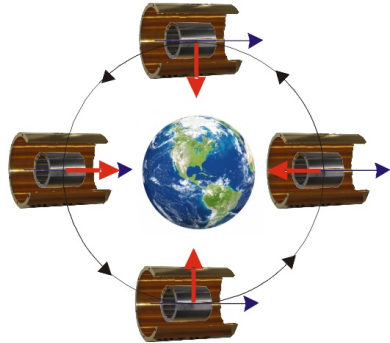
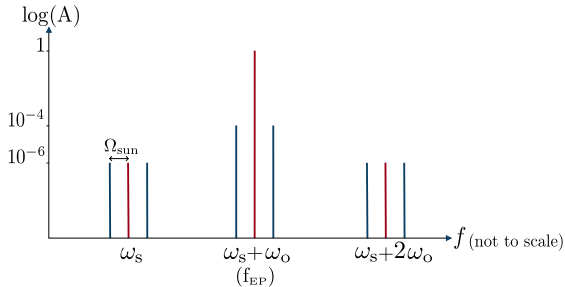


Signal $\neq 0$

Signal $\neq 0$

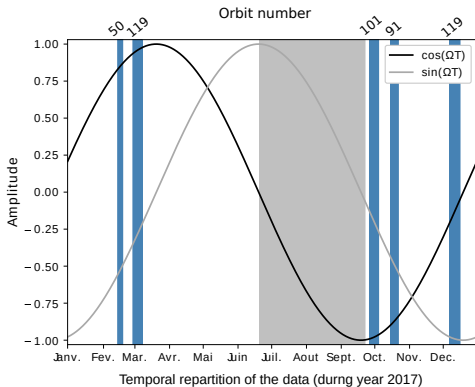
- Test in the sector of couplings between matter and gravitation. Constraints on $\alpha(\bar{a}_{\text{eff}})_{\mu}^W$ components.
- $\alpha(\bar{a}_{\text{eff}})_{\mu}^W$ quantifies a modification massive bodies trajectories depending on their respective composition.

What sort of signals can we expect?

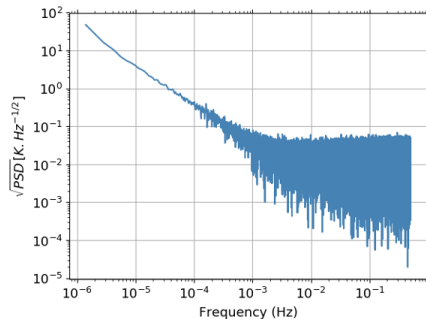
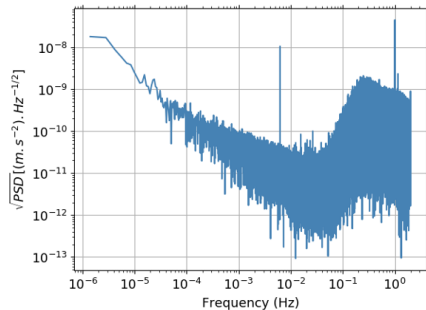


- Variation of the relative acceleration depending on the orientation of the sensitive axis
- Annual modulation due to the motion of the satellite with respect to the Sun.

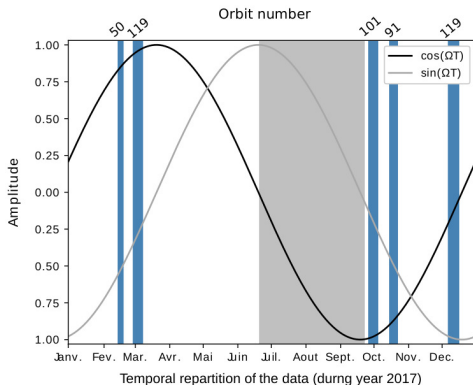
Data used for the analysis



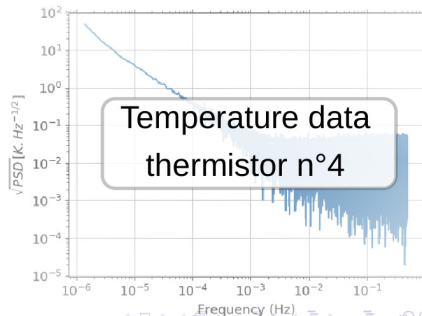
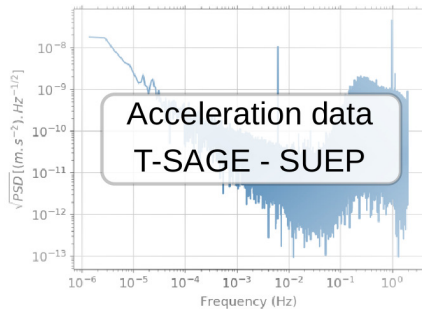
5 sessions - 480 orbits
30 days of cumulated data
gap $< 10^{-3} \%$



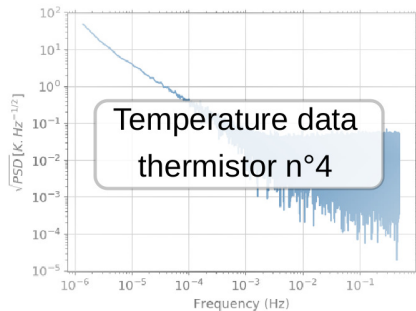
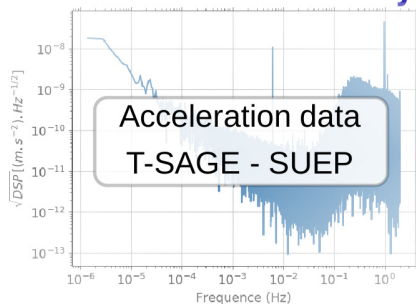
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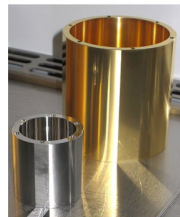
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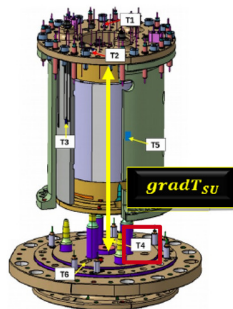


Platinum



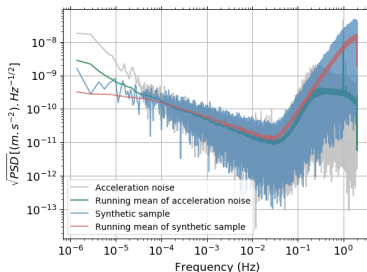
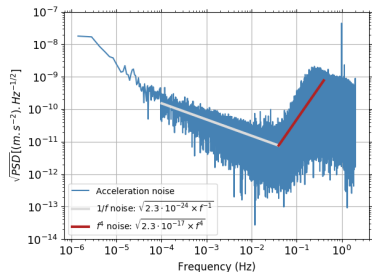
Titanium

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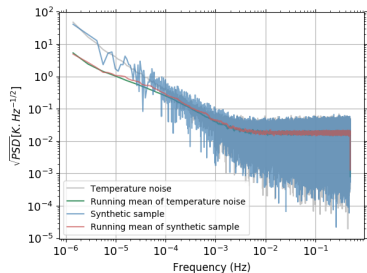
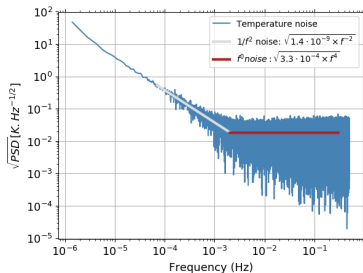
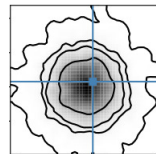


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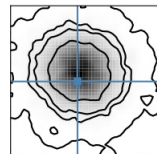
Analysis method: Least-squares Monte Carlo (LSMC) method



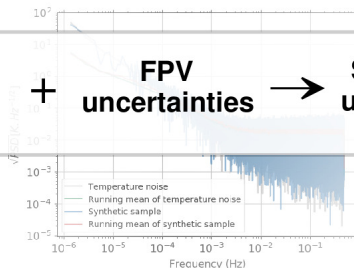
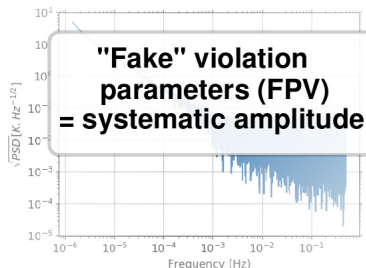
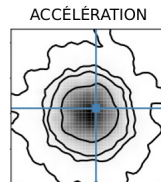
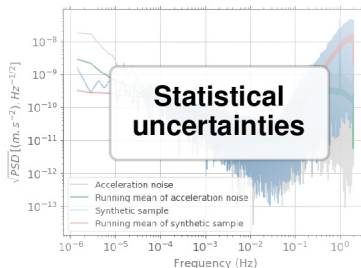
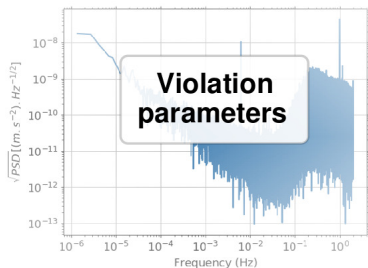
ACCÉLÉRATION



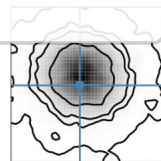
TEMPERATURE



Analysis method: Least-squares Monte Carlo (LSMC) method



Systematic uncertainties



Validation of the method

LSMC vs optimal General Least-squares (GLS)

- Simulated test-data (1000 data points) with realistic noise
- LSMC performances were about a factor 1.4 worse than GLS

WEP analysis with LSMC vs *Touboul et al., PRL, 119, 231101 (2017)*

- LSMC leads to conservative estimation of parameters bounds
- Overall factor 2 between LSMC and the analysis realized in Touboul2017

Parameter	Value and uncertainties		unit
	This analysis	Touboul 2017	
δ	$(-1.3 \pm 2.2 \pm 2.2) \times 10^{-14}$	$(-0.1 \pm 0.9 \pm 0.9) \times 10^{-14}$	—

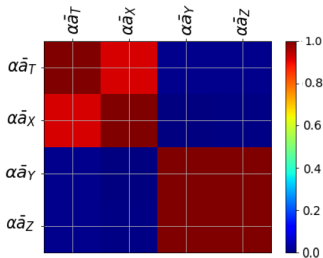
Room for improvement...

Résultats of the SME combined analysis

$$\alpha(a_{\text{eff}}^{(d)})_{\mu} = A \alpha(a_{\text{eff}}^{(n)})_{\mu} + B \alpha(a_{\text{eff}}^{(e+p)})_{\mu}$$

with $A = 0.47 \text{ GeV}^{-1}$ and $B = 0.29 \text{ GeV}^{-1}$

Parameters	Value and uncertainties	Unit	Maximal sensitivities			Unit
			e	p	n	
$\alpha(a_{\text{eff}}^{(d)})_T$	$(0.4 \pm 1.5) (1.1)(1.0) \times 10^{-14}$	—	$10^{-13(2)}$	$10^{-13(2)}$	$10^{-13(2)}$	GeV
$\alpha(a_{\text{eff}}^{(d)})_X$	$(0.6 \pm 2.2) (1.3)(1.8) \times 10^{-10}$	—	$10^{-9(3)}$	$10^{-9(3)}$	$10^{-9(4)}$	GeV
$\alpha(a_{\text{eff}}^{(d)})_Y$	$(0.3 \pm 2.5) (1.1)(2.2) \times 10^{-8}$	—	$10^{-7(2)}$	$10^{-7(2)}$	$10^{-7(2)}$	GeV
$\alpha(a_{\text{eff}}^{(d)})_Z$	$(-0.7 \pm 5.8) (2.5)(5.2) \times 10^{-8}$	—	$10^{-6(1)}$	$10^{-6(1)}$	$10^{-7(2)}$	GeV



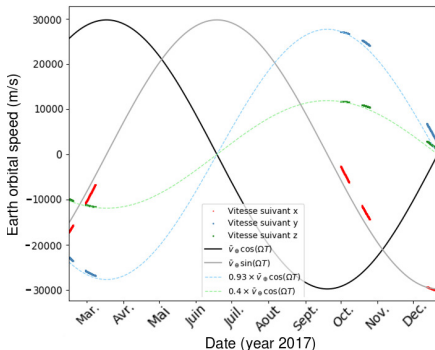
- Combined analysis of 5 sessions
- **No signatures of Lorentz violation detected**
- Improvement by 1 to 4 orders of magnitude on the limites of $\alpha(\bar{a}_{\text{eff}})^w_{\mu}$

This work has been carried out in collaboration with the teams of OCA and ONERA involved in the MICROSCOPE mission, in particular:

G. Métris (OCA), J. Bergé (ONERA) and M. Rodrigues (ONERA)

Thank you for your attention

Correlations in SME analysis



$\beta_X \sim \text{constant} > T$ and X components correlated

β_Y & β_Z vary identically $> Y$ and Z 100 % correlated

$$\begin{aligned}
 2\gamma_{app,\hat{x}}^{(d)} = & S_{\hat{x}\hat{x}}\Delta_{\hat{x}} + (S_{\hat{x}\hat{y}} + \dot{\Omega}_z)\Delta_{\hat{y}} + (S_{\hat{x}\hat{z}} - \dot{\Omega}_y)\Delta_{\hat{z}} \\
 & + g_{\hat{x}} \left[2\alpha(a_{\text{eff}}^{(d)})_T + \beta_X 2\alpha(a_{\text{eff}}^{(d)})_X + \beta_Y 2\alpha(a_{\text{eff}}^{(d)})_Y + \beta_Z 2\alpha(a_{\text{eff}}^{(d)})_Z \right] \\
 & - \frac{3Gm^s R_{\oplus}^2}{5cr^5} \left(R_{\hat{x}\hat{x}}\tilde{x}^{\text{orb}} + R_{\hat{x}\hat{y}}\tilde{y}^{\text{orb}} + R_{\hat{x}\hat{z}}\tilde{z}^{\text{orb}} \right) \left(\tilde{x}^{\text{orb}} 2\alpha(a_{\text{eff}}^{(d)})_Y - \tilde{y}^{\text{orb}} 2\alpha(a_{\text{eff}}^{(d)})_X \right) \omega_{\tilde{z}} \\
 & + \frac{Gm^s R_{\oplus}^2}{5cr^3} \left(2\alpha(a_{\text{eff}}^{(d)})_Y R_{\hat{x}\hat{x}} - 2\alpha(a_{\text{eff}}^{(d)})_X R_{\hat{x}\hat{y}} \right) \omega_{\tilde{z}}
 \end{aligned} \quad (2.43)$$

Mission data

INPOP

Estimated in
our analysis