

Transportable optical lattice clocks



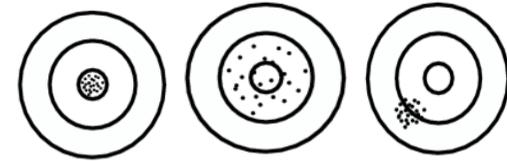
Christian Lisdat

ESA Topical Team Meeting, 22. – 23.10.2018, Munich

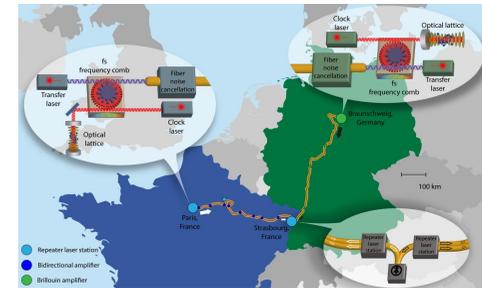


Outline:

► Few words about (optical) clocks



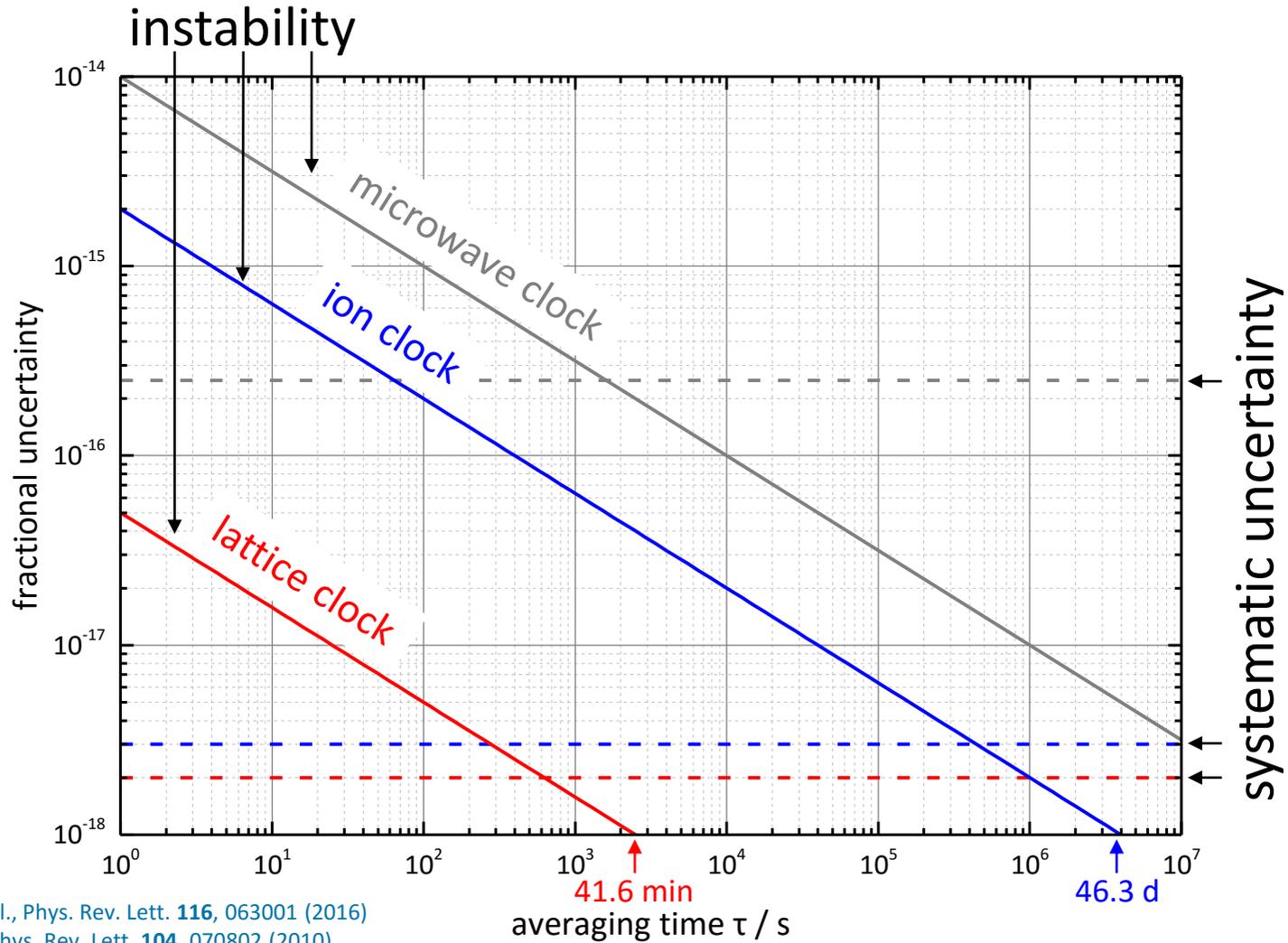
► Clocks in the lab: examples of experiments



► Campaigns with our transportable optical lattice clock



Best clock performances



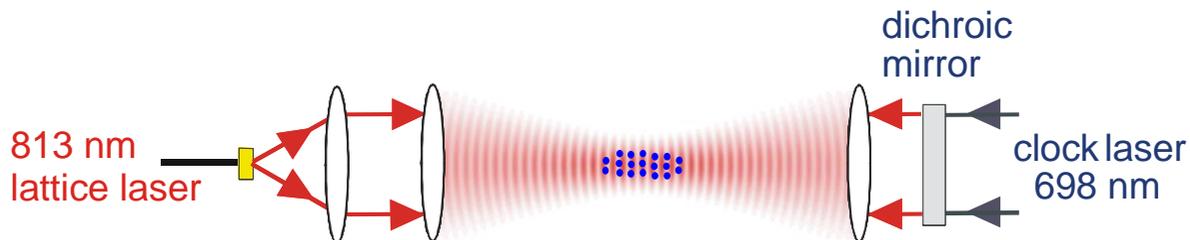
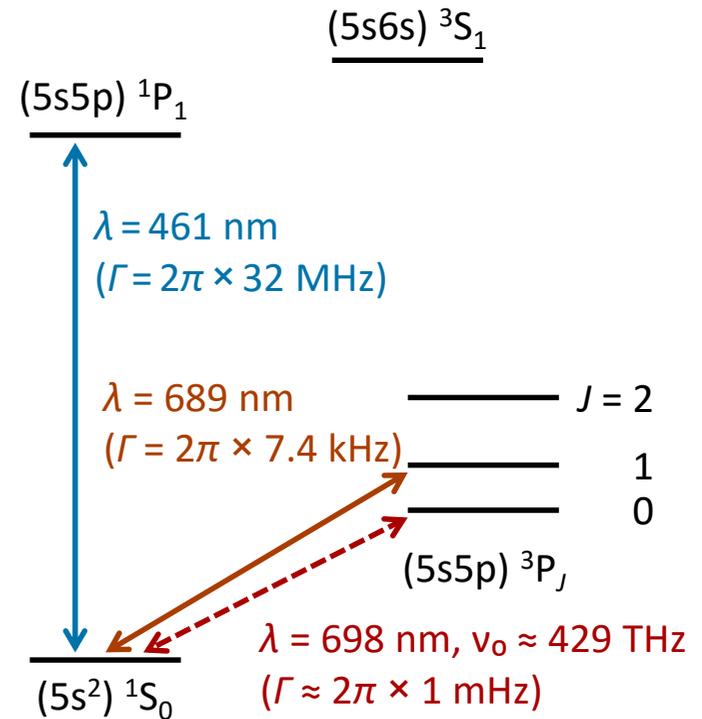
- N. Huntemann et al., Phys. Rev. Lett. **116**, 063001 (2016)
- C. W. Chou et al., Phys. Rev. Lett. **104**, 070802 (2010)
- T. Nicholson et al., Nature Com. **6**, 6896 (2015)
- M. Schioppo et al., Nature Photonics **11**, 48 (2017)

Strontium lattice clock

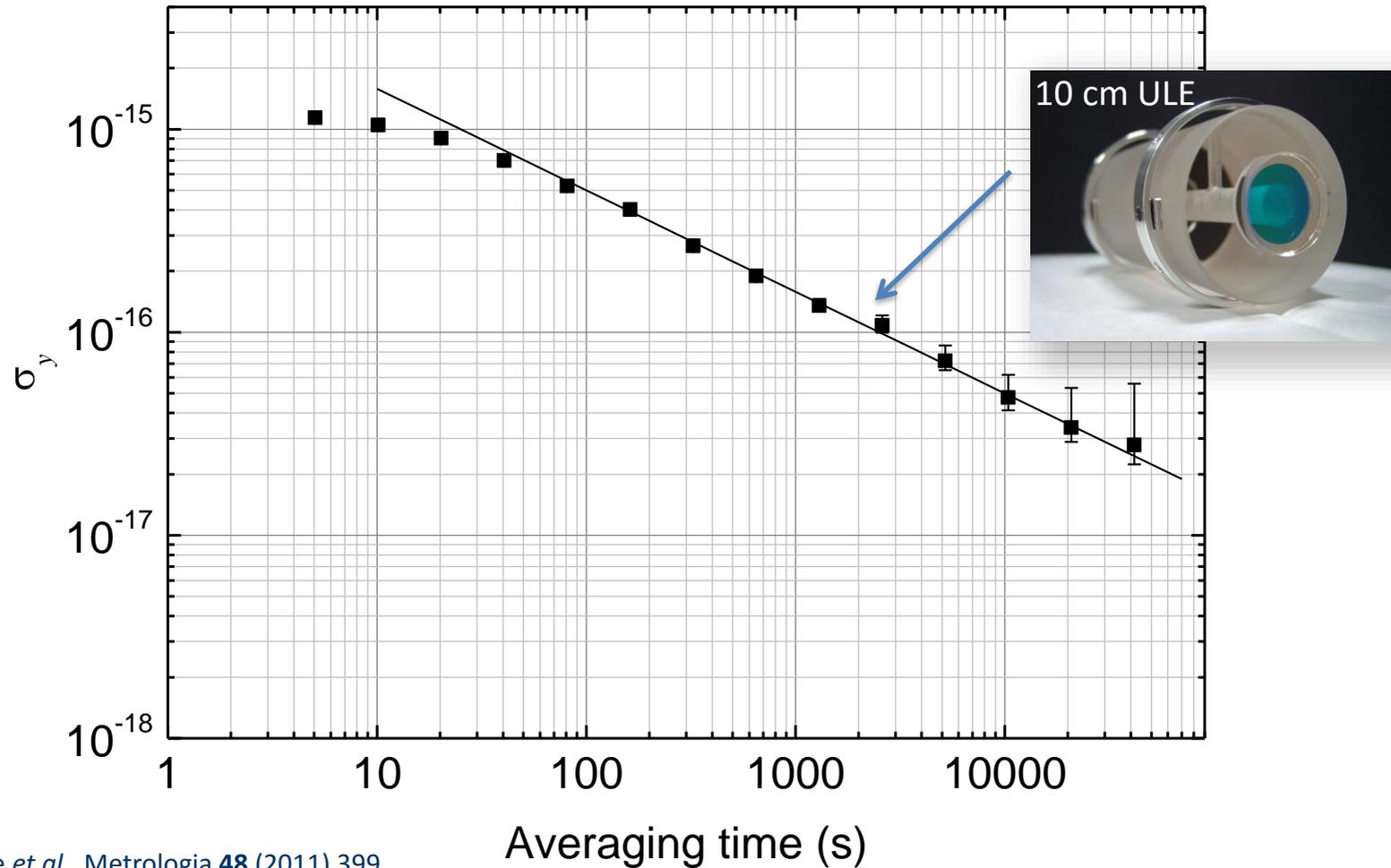
Experimental sequence

200 ms	Zeeman slower 1 st stage MOT (461 nm), $T \sim$ mK
90 ms	2 nd stage MOT (689 nm)
90 ms	3 rd stage MOT (689 nm), $T \sim$ μ K
65 ms	state preparation in 1D optical lattice
0.8 s – 2.6 s	Rabi interrogation (698 nm) in 1D optical lattice
	balanced detection

Partial level diagram

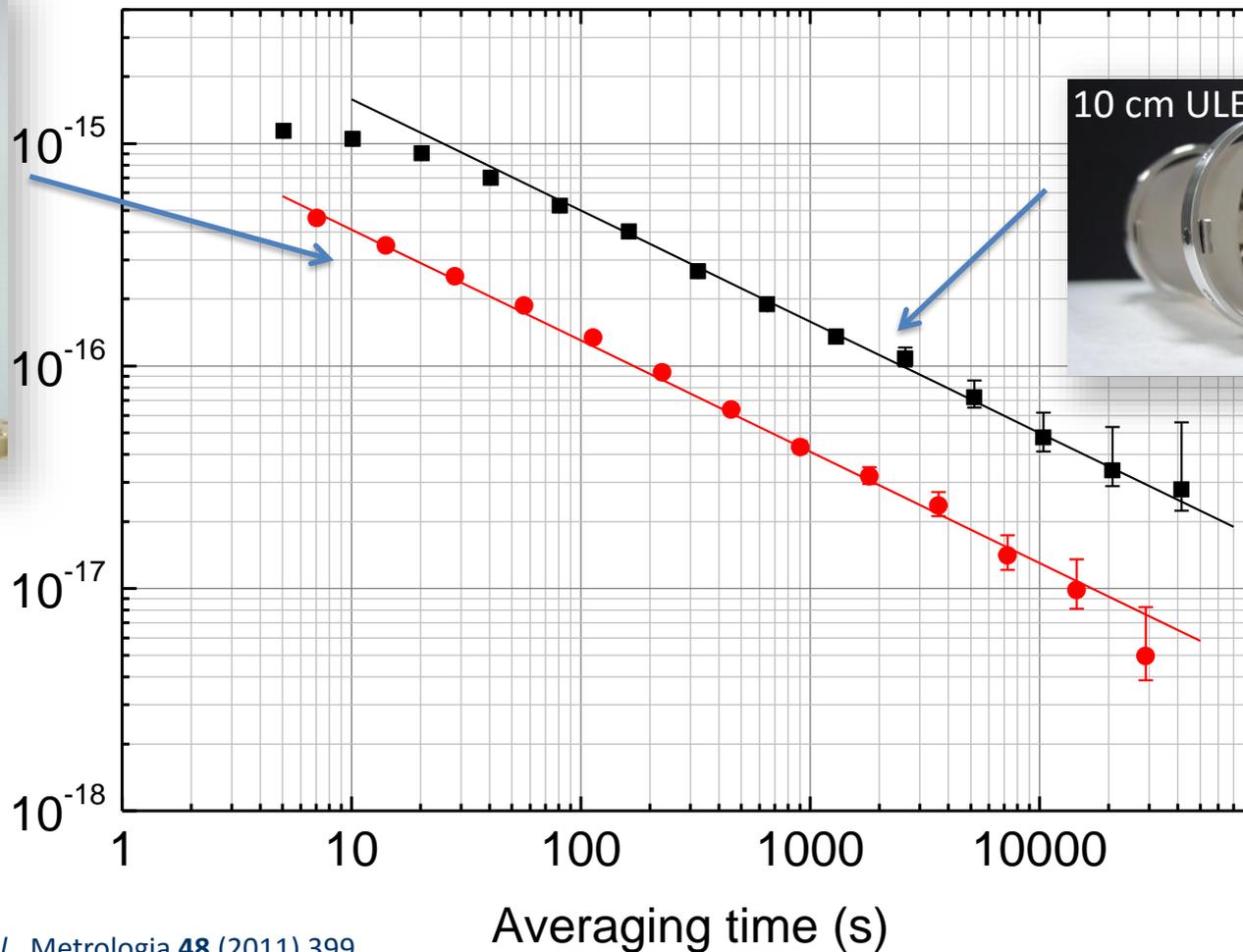


Importance of the clock laser



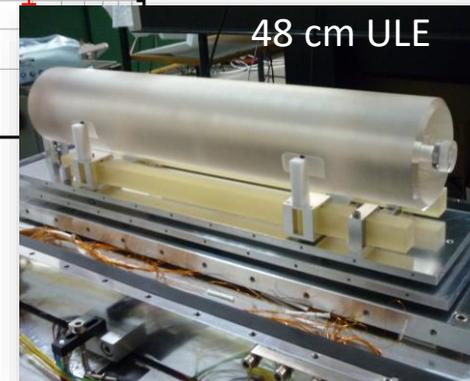
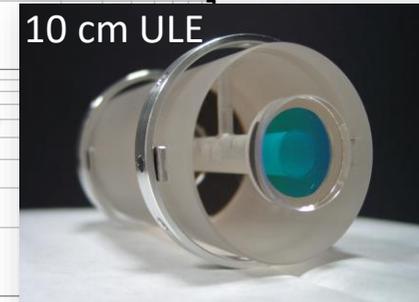
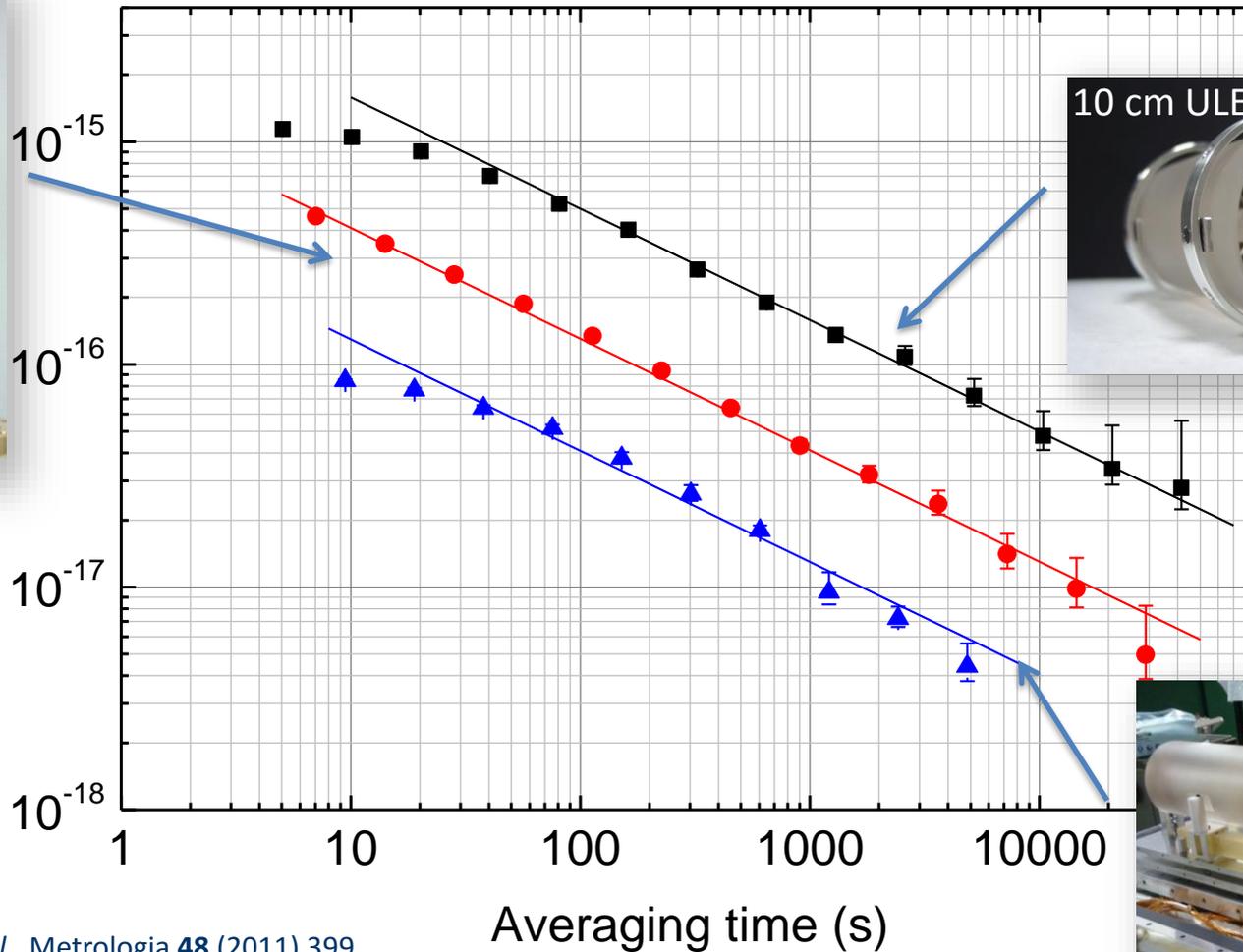
St. Falke *et al.*, *Metrologia* **48** (2011) 399 ,
Ch. Hagemann *et al.*, *IEEE Trans. Instr. Meas.* **62** (2013) 1556
A. Al-Masoudi *et al.*, *PRA* **92**, 063814 (2015)

Importance of the clock laser



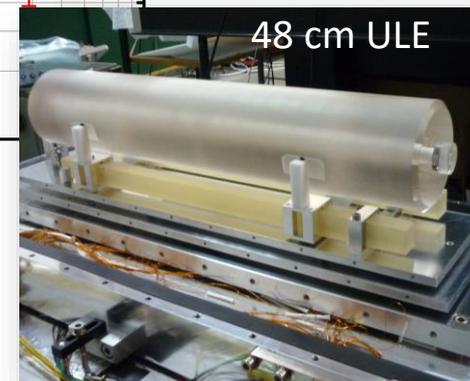
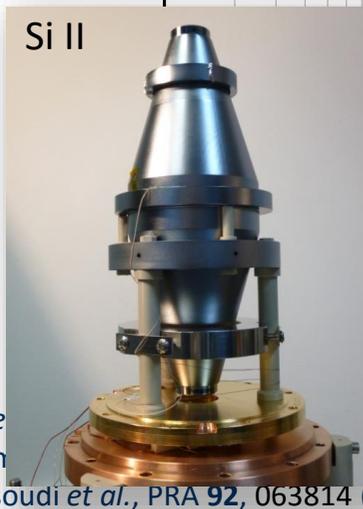
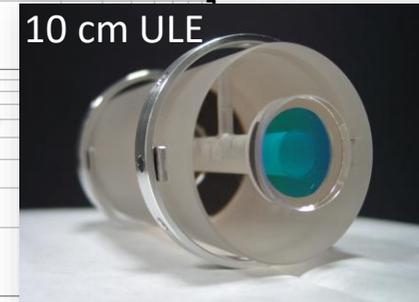
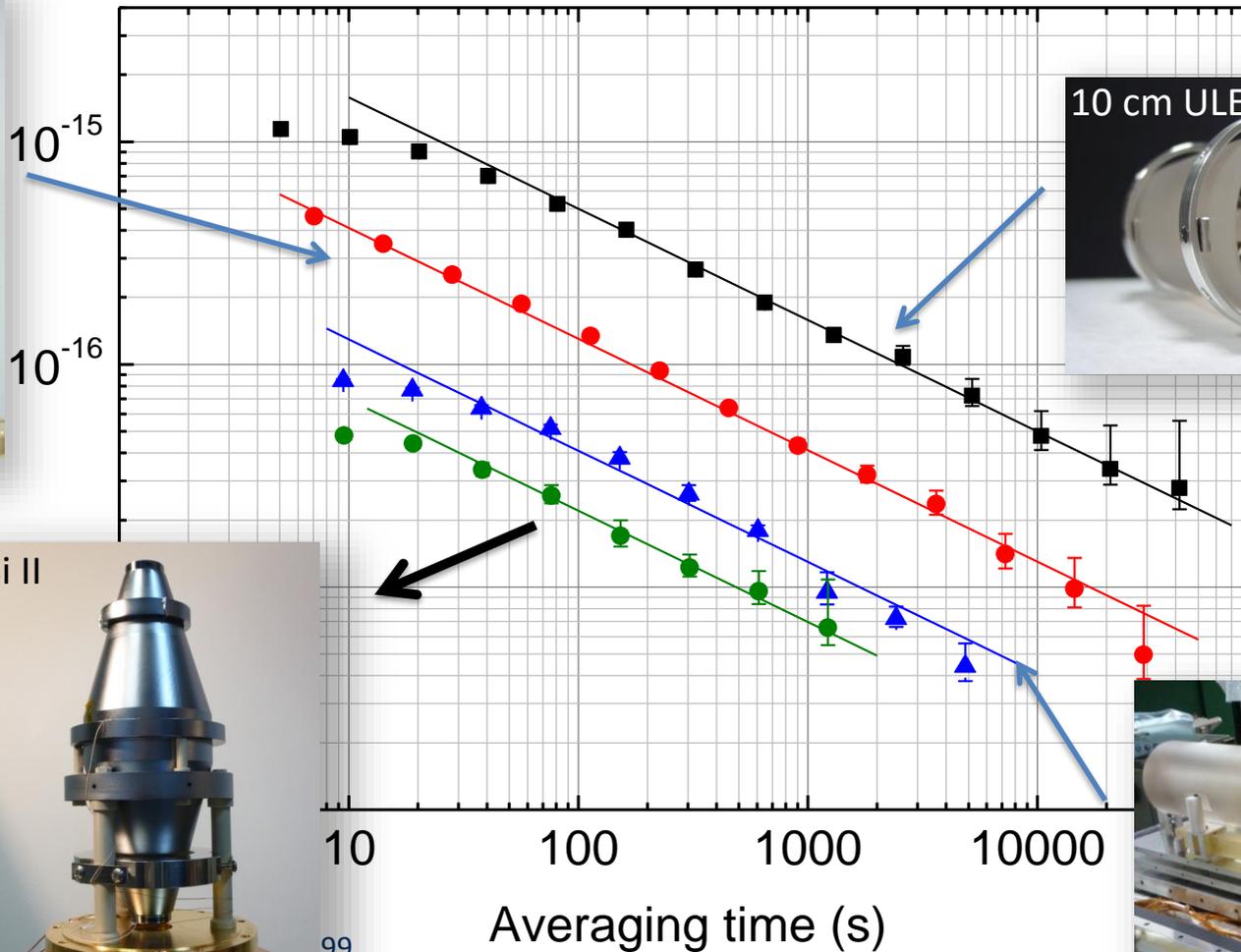
St. Falke *et al.*, *Metrologia* **48** (2011) 399 ,
Ch. Hagemann *et al.*, *IEEE Trans. Instr. Meas.* **62** (2013) 1556
A. Al-Masoudi *et al.*, *PRA* **92**, 063814 (2015)

Importance of the clock laser



St. Falke *et al.*, *Metrologia* **48** (2011) 399 ,
Ch. Hagemann *et al.*, *IEEE Trans. Instr. Meas.* **62** (2013) 1556
A. Al-Masoudi *et al.*, *PRA* **92**, 063814 (2015)

Importance of the clock laser



St. Falke et al., Phys. Rev. Lett. **99**, 150801 (2002)
 Ch. Hagen et al., Phys. Rev. Lett. **99**, 150802 (2002)
 A. Al-Masoudi et al., Phys. Rev. A **92**, 063814 (2015)

Importance of the clock laser



FEM optimized cavity shape for minimal vibration sensitivity

expected thermal noise limit at $T = 123.5 \text{ K}$:

$$\text{mod } \sigma_y \approx 4 \times 10^{-17}$$

$$\frac{\Delta\nu}{\nu} = -\frac{\Delta L}{L}$$

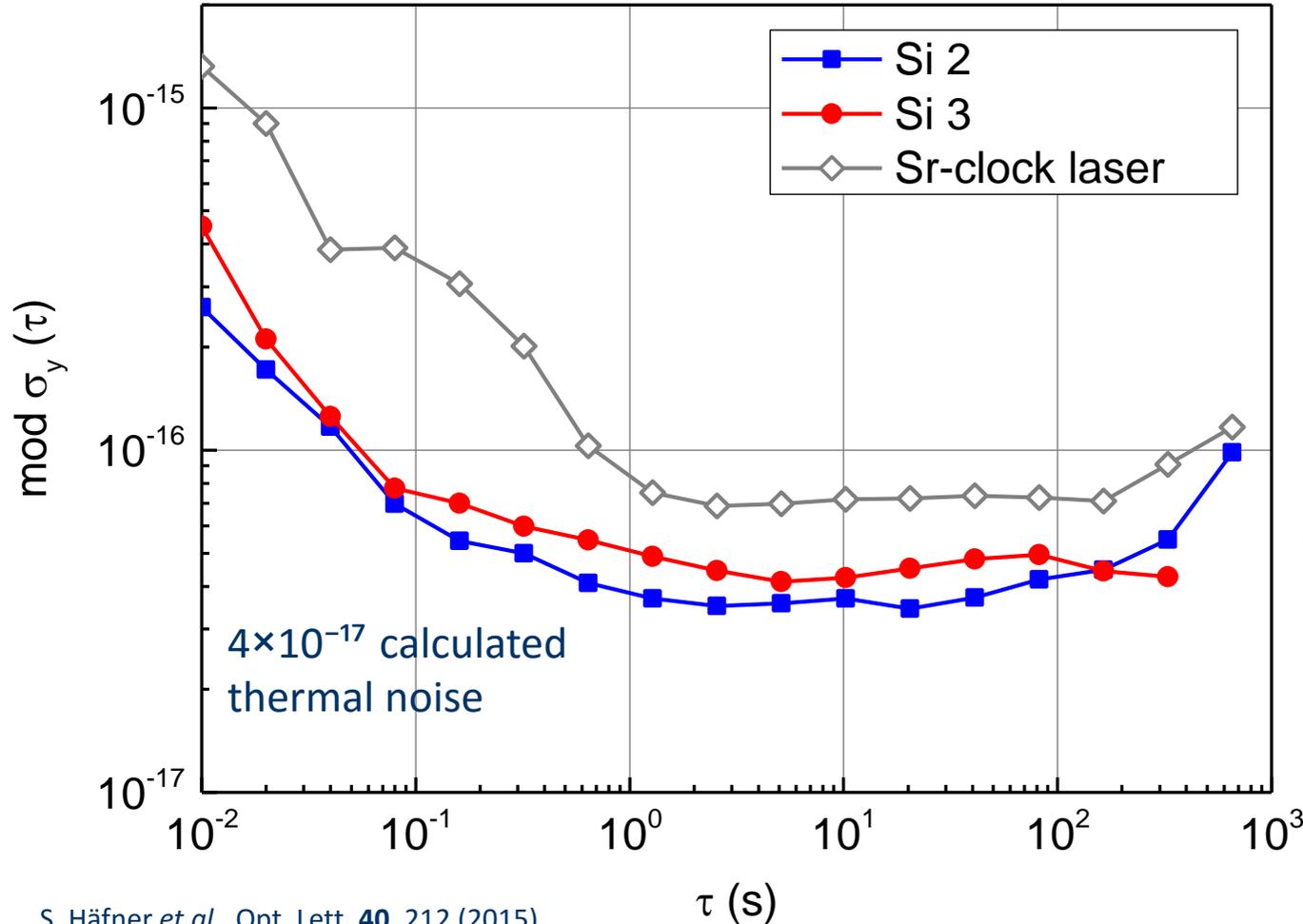
absolute length fluctuations
 $\approx 8.5 \times 10^{-18} \text{ m}$

dominated by mirror coatings!

proton diameter $\approx 0.85 \text{ fm} = 850 \times 10^{-18} \text{ m}$

T. Kessler *et al.*, Nature Phot. **6**, 687 (2012),
D. Matei *et al.*, Phys. Rev. Lett. **118**, 263202 (2017)

Importance of the clock laser



linear drift removed
from Si 3 – Si 2
and Sr – Si 2 beats

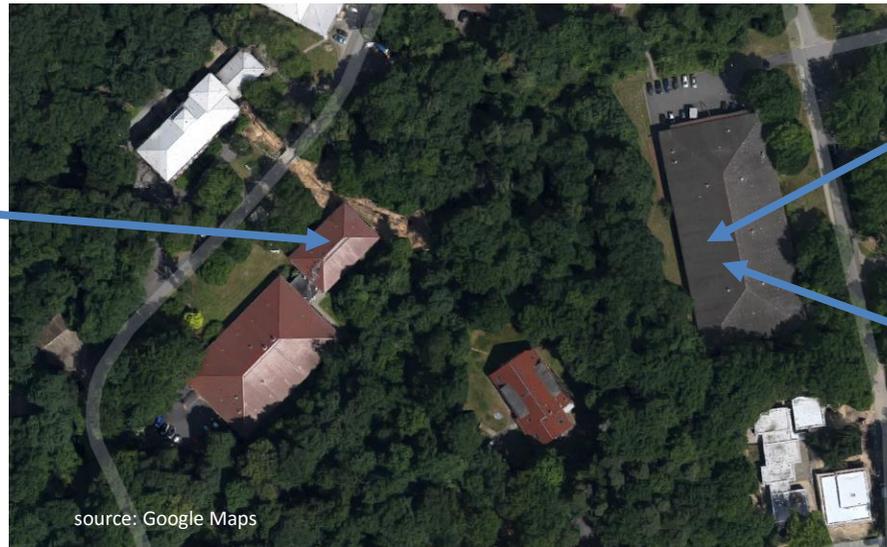
7×10^{-17} thermal noise
of ULE cavity

S. Häfner *et al.*, Opt. Lett. **40**, 212 (2015),
T. Kessler *et al.*, Nature Phot. **6**, 687 (2012),
D. Matei *et al.*, Phys. Rev. Lett. **118**, 263202 (2017)

What can you do?



Yb⁺ single ion clock



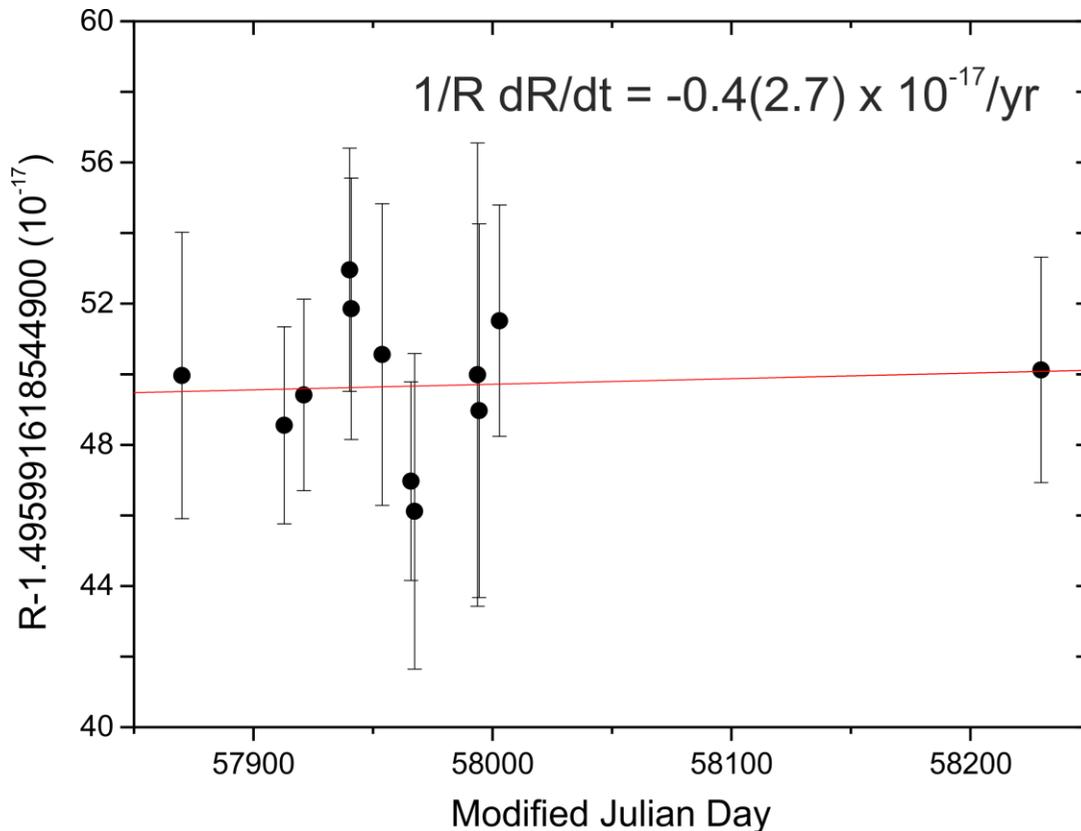
building A

building B



“Space Optical Sr Clock”

Clock comparisons – local and non-local



$$\mathcal{R} = \nu(\text{Yb}^+, E3) / \nu(\text{Sr})$$

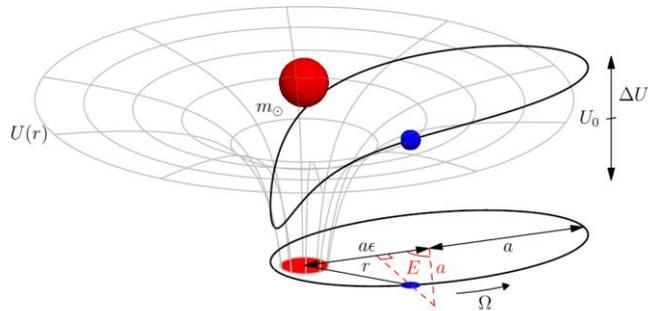
$$\frac{\Delta \mathcal{R}}{\mathcal{R}} = -6 \frac{\Delta \alpha}{\alpha}$$

- Yb⁺/Sr frequency ratio magnifies α changes by 6
- Systematic uncertainty of the Sr clock 2×10^{-17}

C. Grebing et al., *Optica* **3**, 563 (2016).

N. Huntemann et al., *PRL* **113**, 210802 (2014)

Repeated comparisons – interpretation

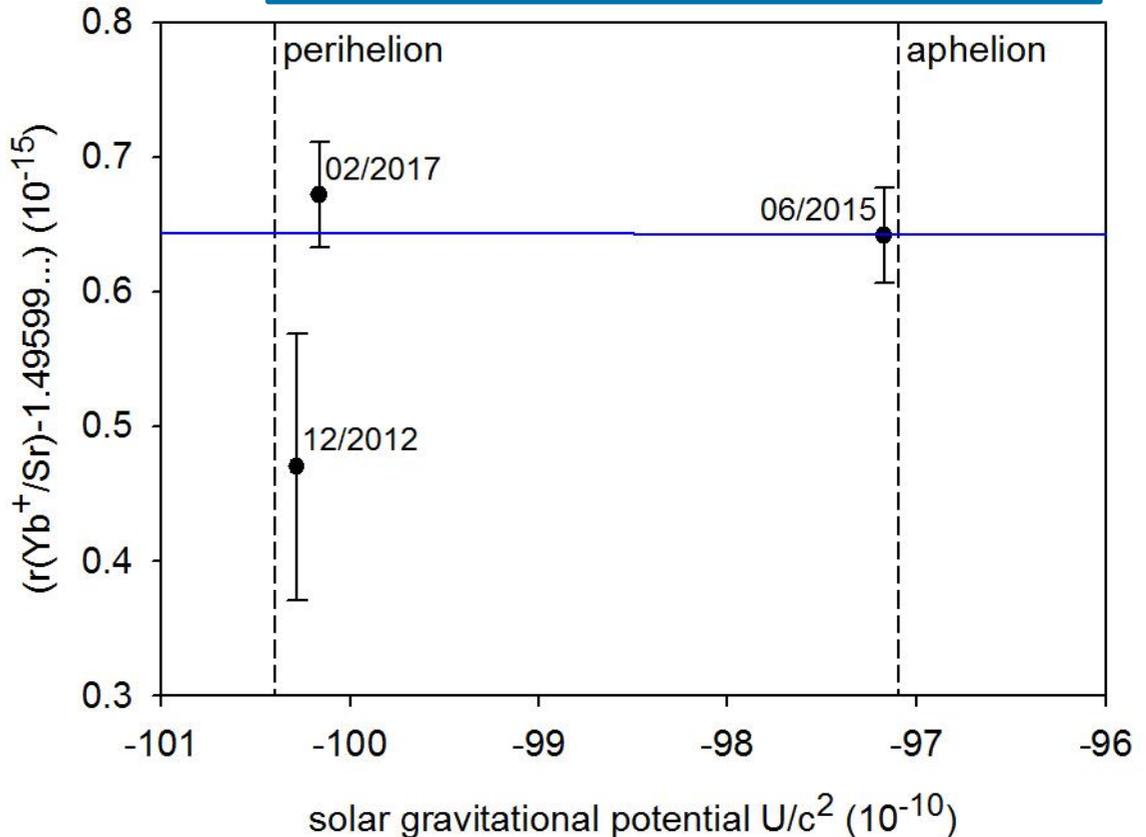


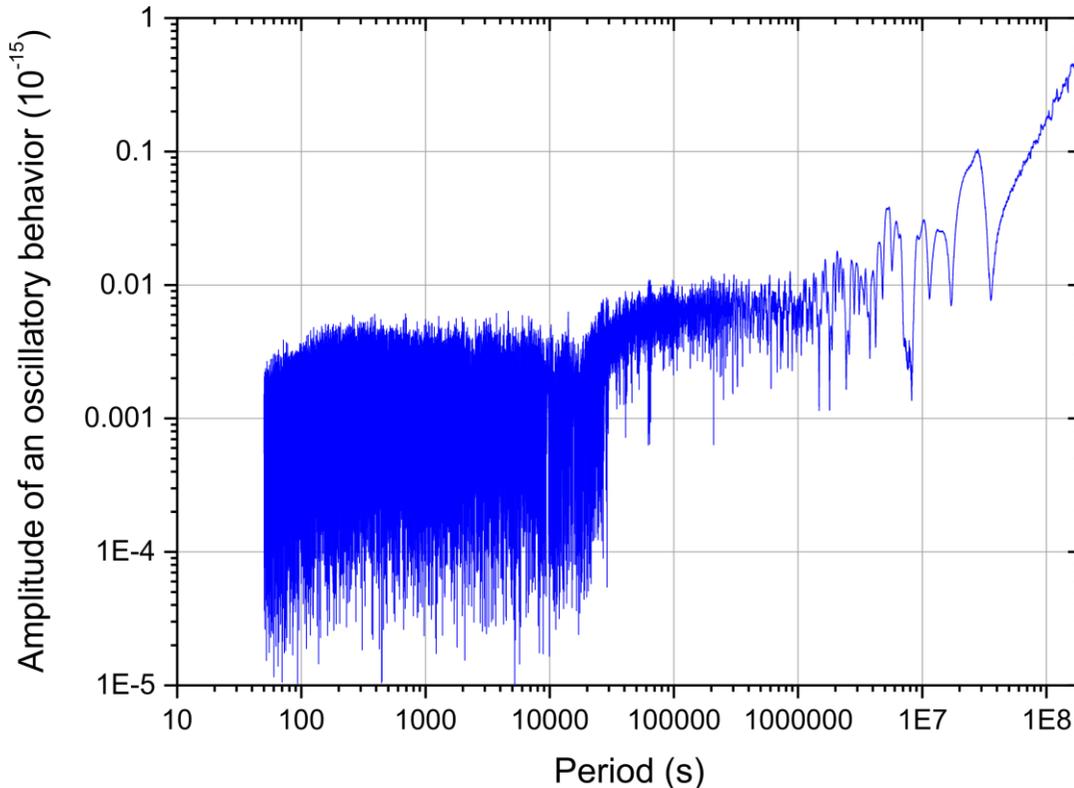
from: S. Blatt et al., Phys. Rev. Lett. **100**, 140801 (2008)

$$(c^2/\alpha) d\alpha/dU = 2.5(3.1) \times 10^{-6}$$

S. Blatt et al., Phys. Rev. Lett. **100**, 140801 (2008)

$$(c^2/\alpha) d\alpha/dU = 0.1(1.9) \times 10^{-8}$$





$$\mathcal{R} = \nu(\text{Yb}^+, E3) / \nu(\text{Sr})$$

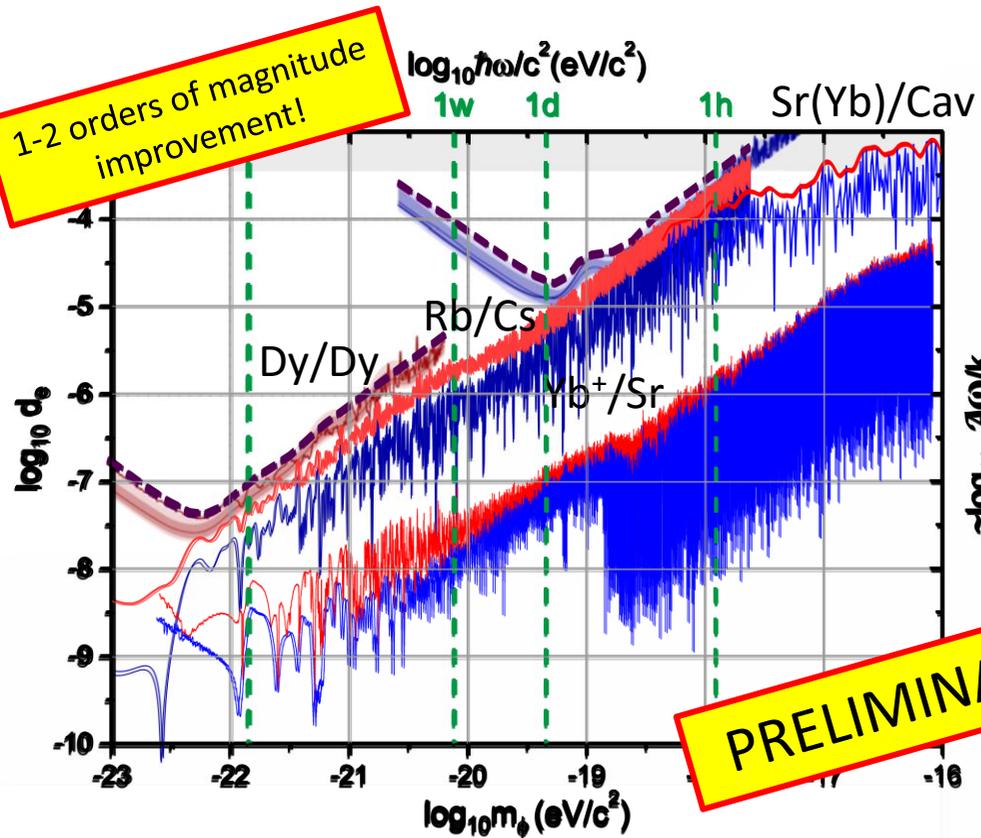
$$\frac{\Delta \mathcal{R}}{\mathcal{R}} = -6 \frac{\Delta \alpha}{\alpha}$$

- Yb⁺/Sr frequency ratio magnifies α changes by 6
- Systematic uncertainty of the Sr clocks 2×10^{-17}
- No oscillatory behaviour of the frequency ratio
- More than 24 days of data

C. Grebing et al., *Optica* **3**, 563 (2016).

N. Huntemann et al., *PRL* **113**, 210802 (2014)

Clock comparisons – local and non-local



Hypothetical coupling of dark matter to the electromagnetic field tensor leads to an oscillation in α and in turn in \mathcal{R} with an amplitude:

$$A = k_\alpha d_e \frac{1}{\omega} \left(\frac{8\pi G}{c^2} \rho_{\text{DM}} \right)^{1/2},$$

$$\rho_{\text{DM}} \approx 0.4 \frac{\text{GeV}}{\text{cm}^3}$$

Ratio measurements provide limits on d_e .

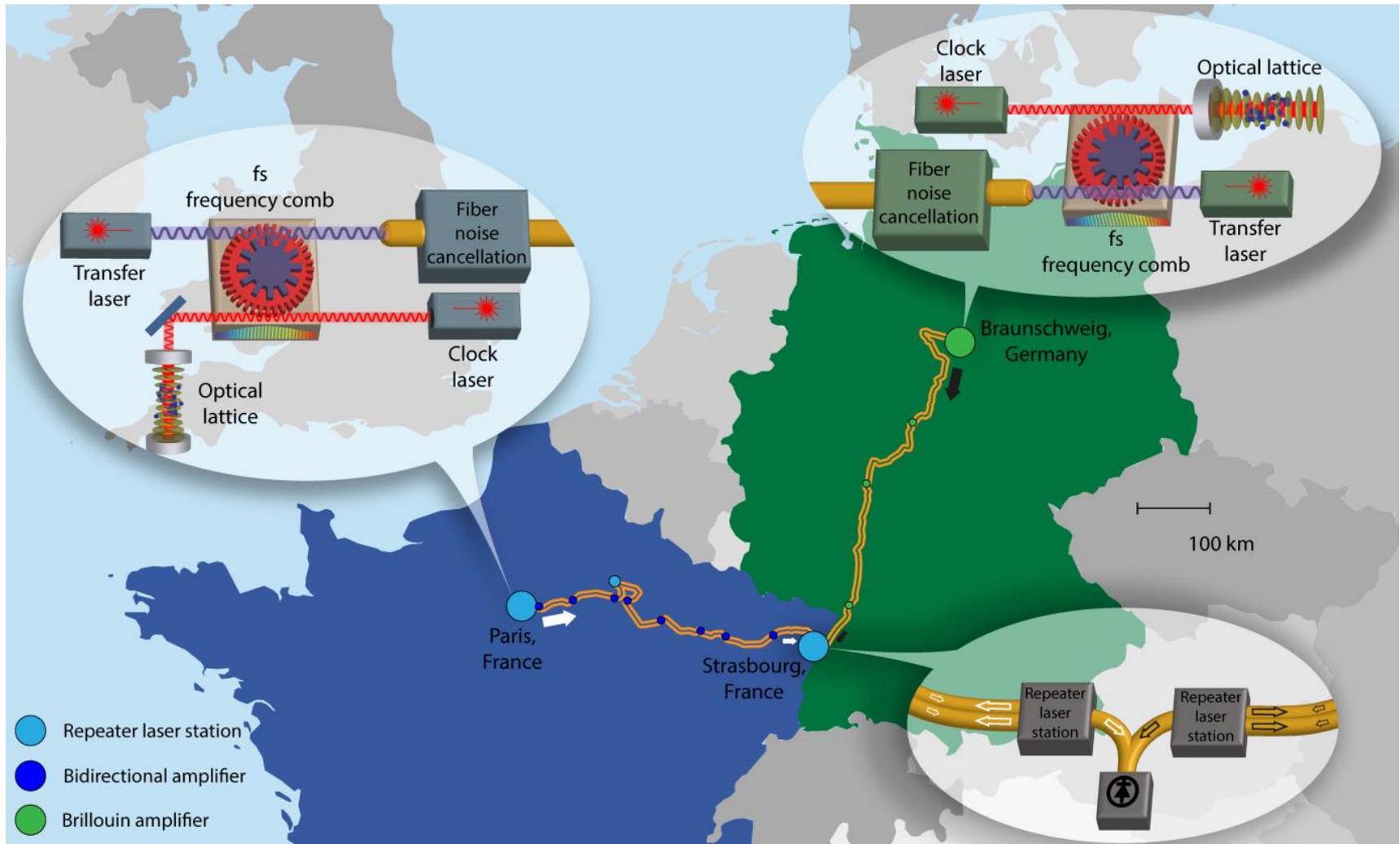
A. Arvanitaki et al., PRD 91, 015015 (2015)

Dy/Dy: K. Van Tilburg et al.,
PRL 115, 011802 (2015)

Rb/Cs: A. Hees et al.,
PRL 117, 061301 (2016)

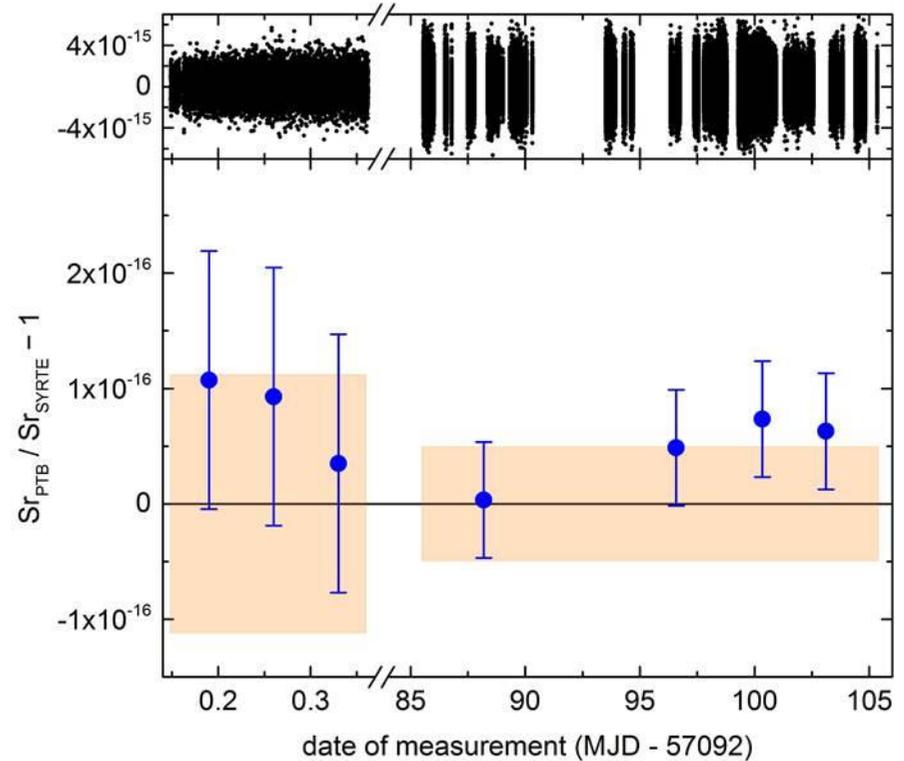
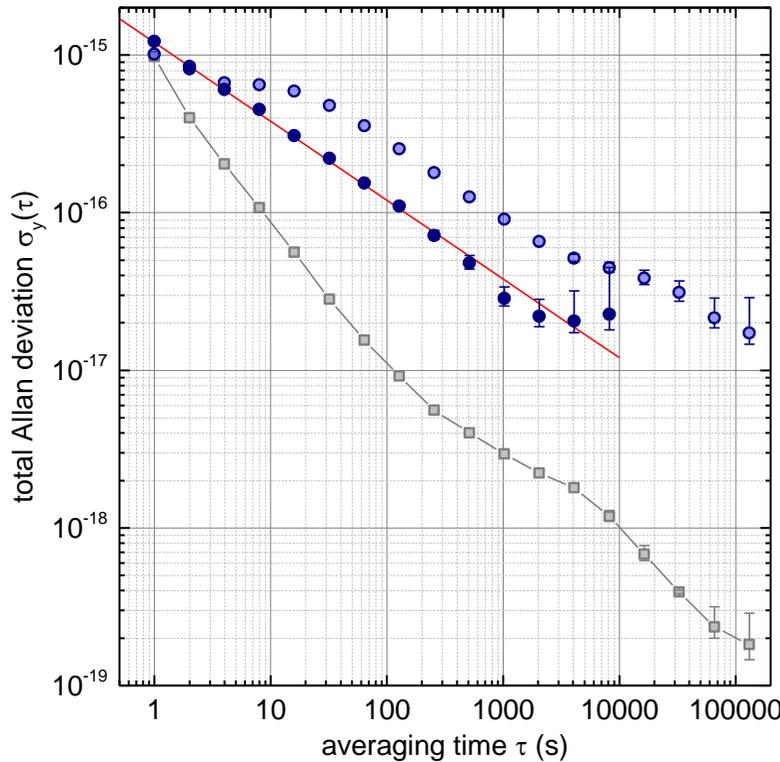
Sr(Yb)/Cav: P. Wcislo et al.,
arXiv:180604762 (2018)

Clock comparisons – local and non-local



Ch. Lisdat *et al.*, Nature Comms. **7**, 12443 (2016)

Clock comparisons – Paris & Braunschweig



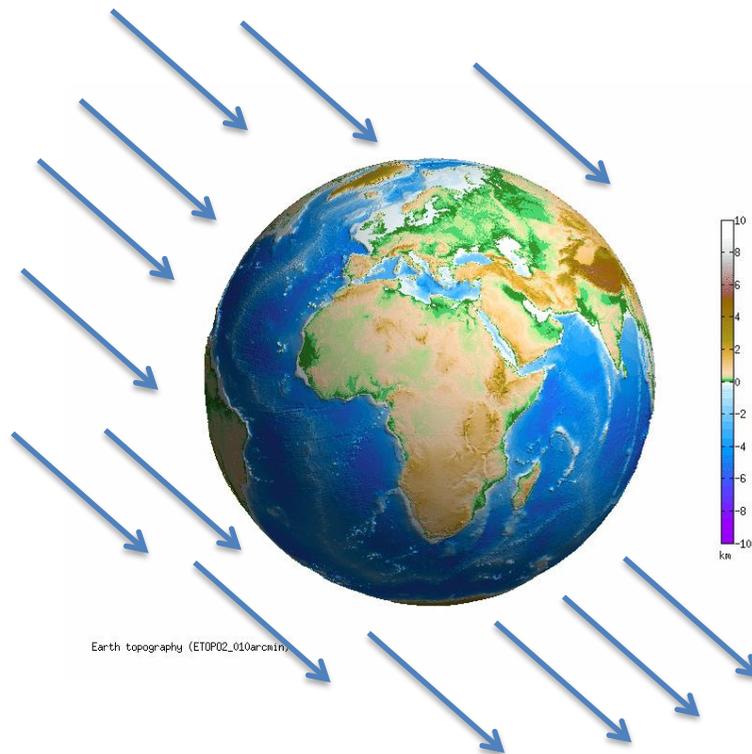
Gravity potential correction
 $-247.2(4) \times 10^{-17}$



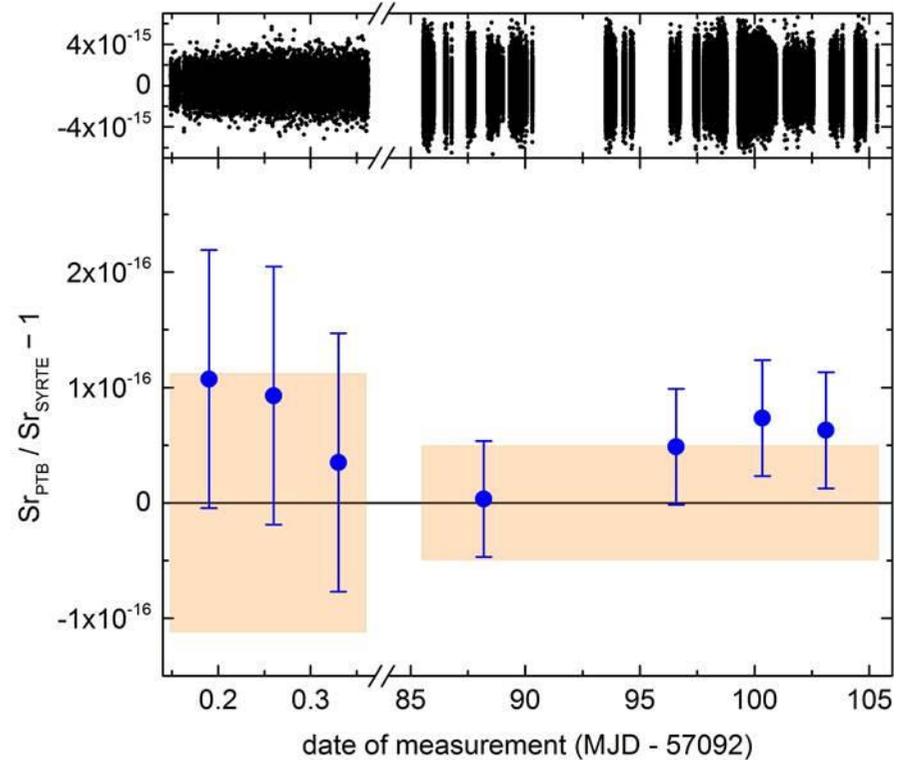
Ch. Lisdat *et al.*, Nature Comms. **7**, 12443 (2016)

Clock comparisons – Paris & Braunschweig

Local Lorentz invariance: search for daily modulation due to motion wrt. background

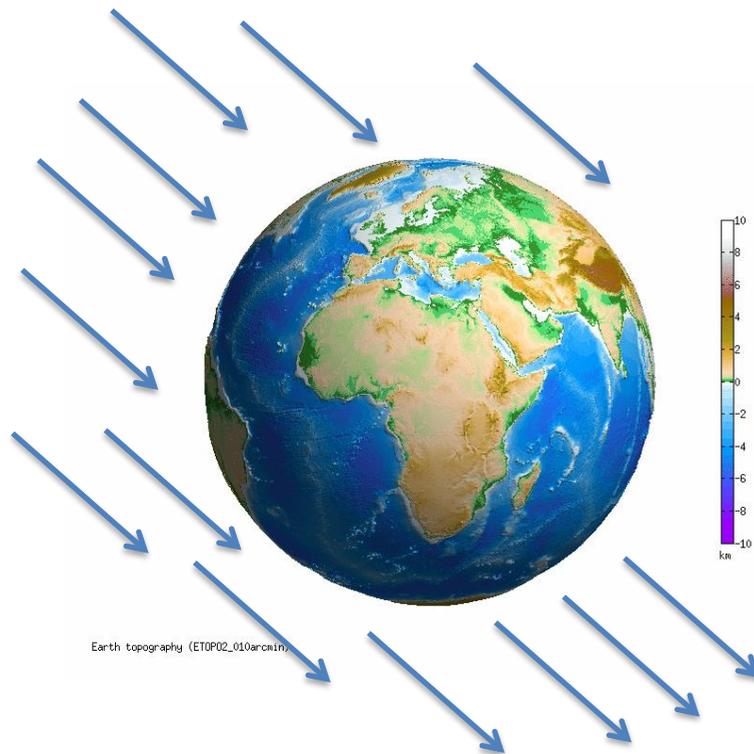


Earth topography (ETOPO2_010arcmin)



animation: A. Bezdek and J. Sebera, Computers & Geosciences **56**, 127 (2013), data set: ETOPO2 / EGM2008

Local Lorentz invariance: search for daily modulation due to motion wrt. background



Earth topography (ETOPO2_010arcmin)

was done with Rb clocks (GPS)
P. Wolf & G. Petit, Phys. Rev. A **56**, 4405 (1997)

$$|\alpha| \leq 10^{-6}$$

LLI test also with fast ion beams
B. Botermann *et al.*, Phys. Rev. Lett. **113**, 120405 (2014)

$$|\alpha| \leq 2 \times 10^{-8}$$

Sr clocks London, Paris, Braunschweig
P. Delva *et al.*, Phys. Rev. Lett. **118**, 221102 (2017)

$$|\alpha| \leq 1.2 \times 10^{-8}$$

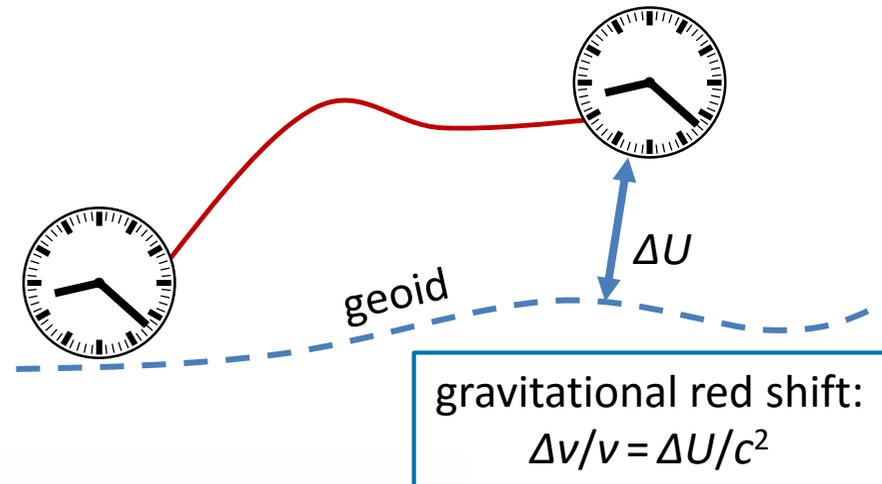
animation: A. Bezdek and J. Sebera, Computers & Geosciences **56**, 127 (2013), data set: ETOPO2 / EGM2008

Leaving the lab: Transportable clocks

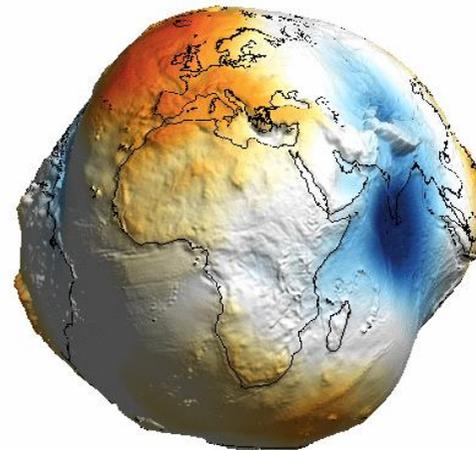
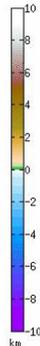
- ▶ flexibility of clock pairs
- ▶ choose operation sites
to probe the gravity potential
- ▶ first step towards space

▶ Optical clocks as sensors:

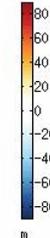
- Directly measure potential differences.
- **Vision:** Realize geoid by clocks.



Earth topography (ETOPO2_010arcmin)

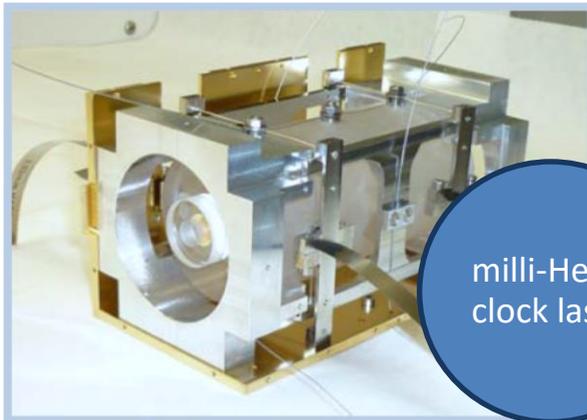


Geoid height (EGM2008, nmax=500)

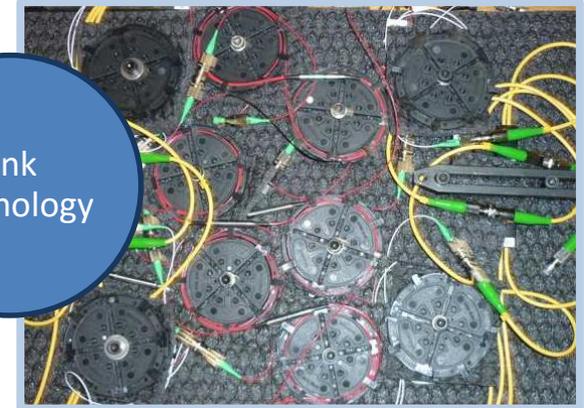


M. Vermeer, Rep. of the Finnish Geod. Insti. **83**, 1 (1983)
 A. Bjerhammar, Bull. Geodesique **59**, 207 (1985)

Transportable optical clocks

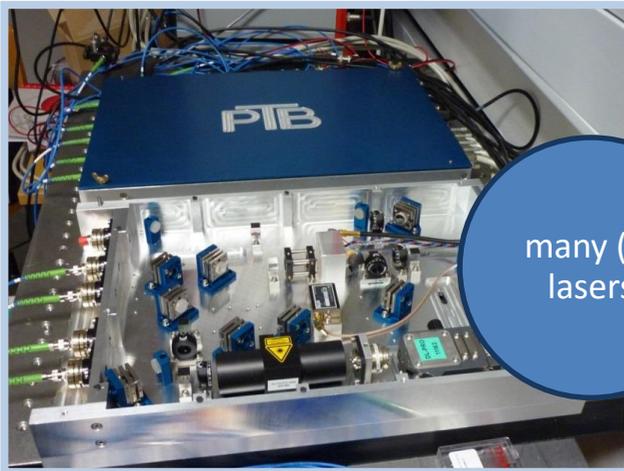


milli-Hertz
clock laser

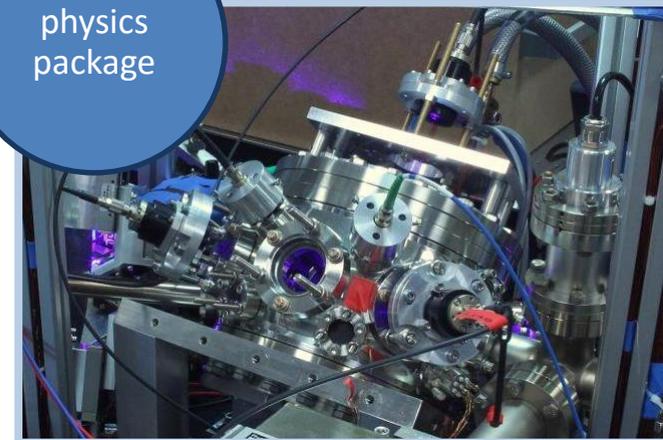


link
technology

Sr



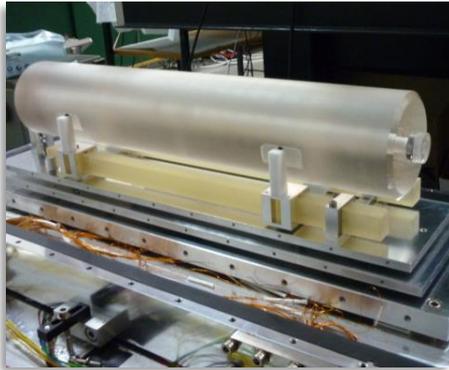
many (6)
lasers



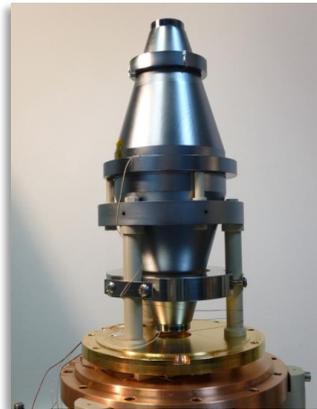
physics
package

Importance of the clock laser

Clock lasers in labs:



48 cm ULE cavity



cryogenic 21 cm silicon cavity

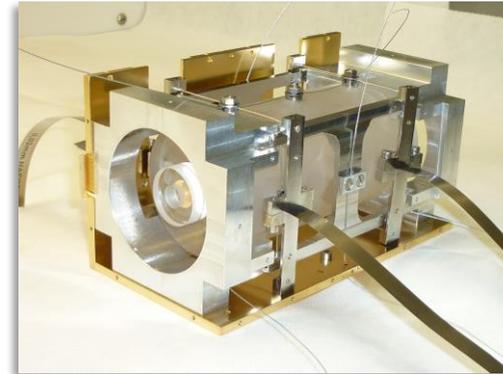
S. Häfner *et al.*, *Opt. Lett.* **40**, 212 (2015),
D. Matei *et al.*, *Phys. Rev. Lett.* **118**, 263202 (2017)

acceleration
sensitivity

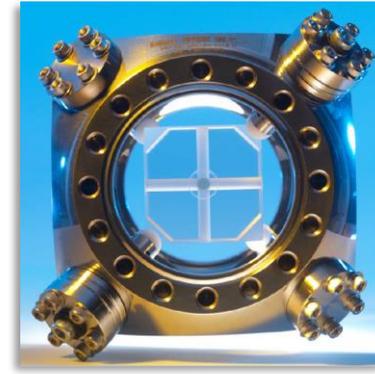


thermal noise

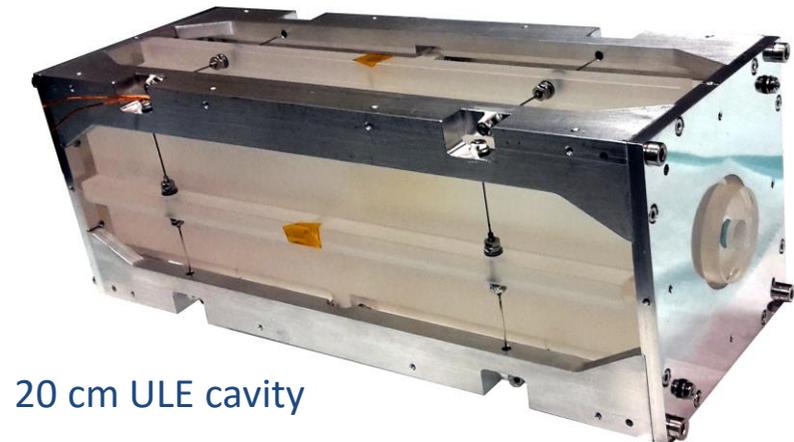
Transportable clock lasers:



12 cm ULE cavity



5 cm ULE cavity



20 cm ULE cavity

S. Koller *et al.*, *Phys. Rev. Lett.* **118**, 073601 (2017)
S. Webster & P. Gill, *Opt. Lett.* **36**, 3572 (2011)

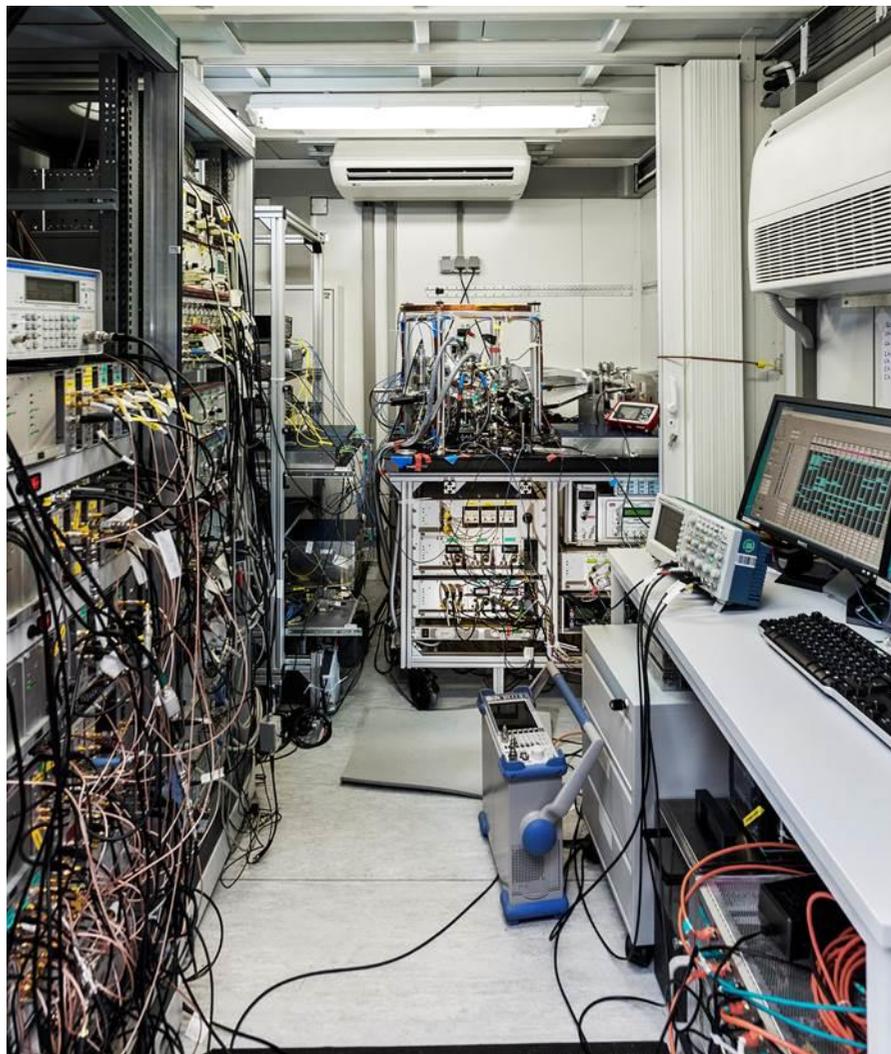
Transportable optical clocks



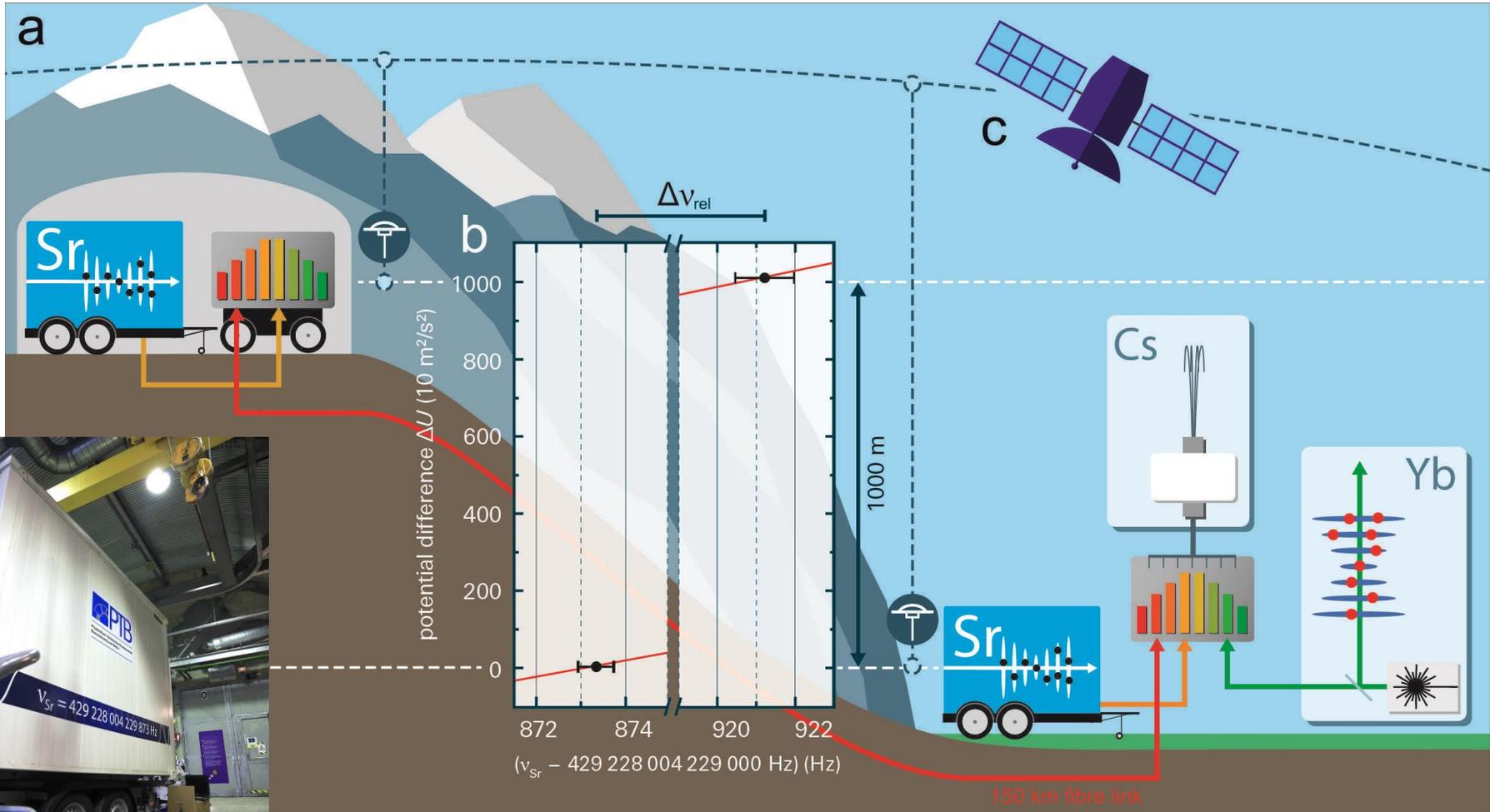
Car trailer housing the clock

S. Koller *et al.*, Phys. Rev. Lett. **118**, 073601 (2017)

View into the car trailer ►

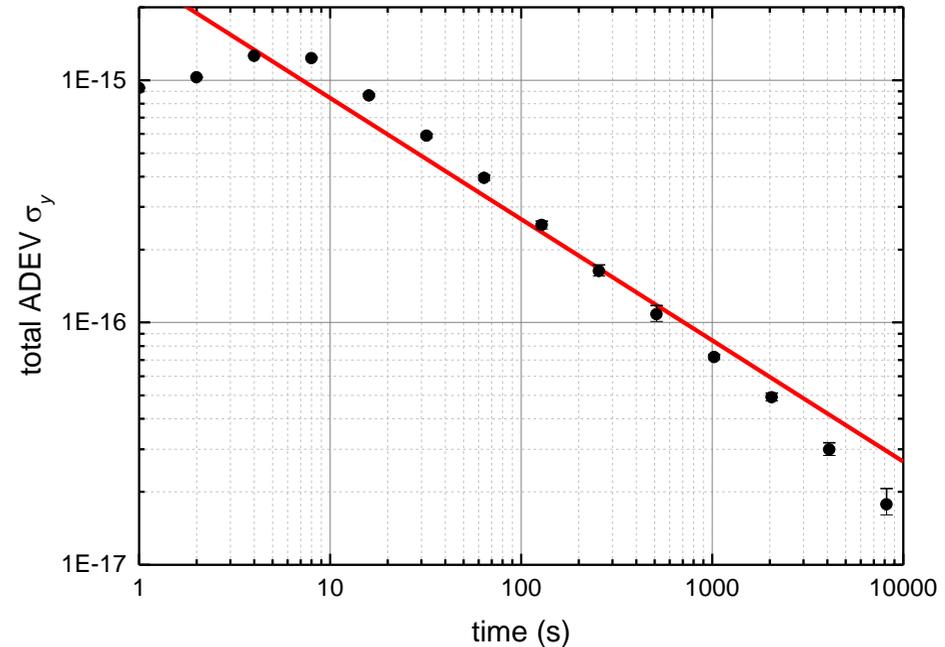


First time off-campus: Modane – Torino 2016



J. Grotti *et al.*, arXiv 1705.04089 (2017)

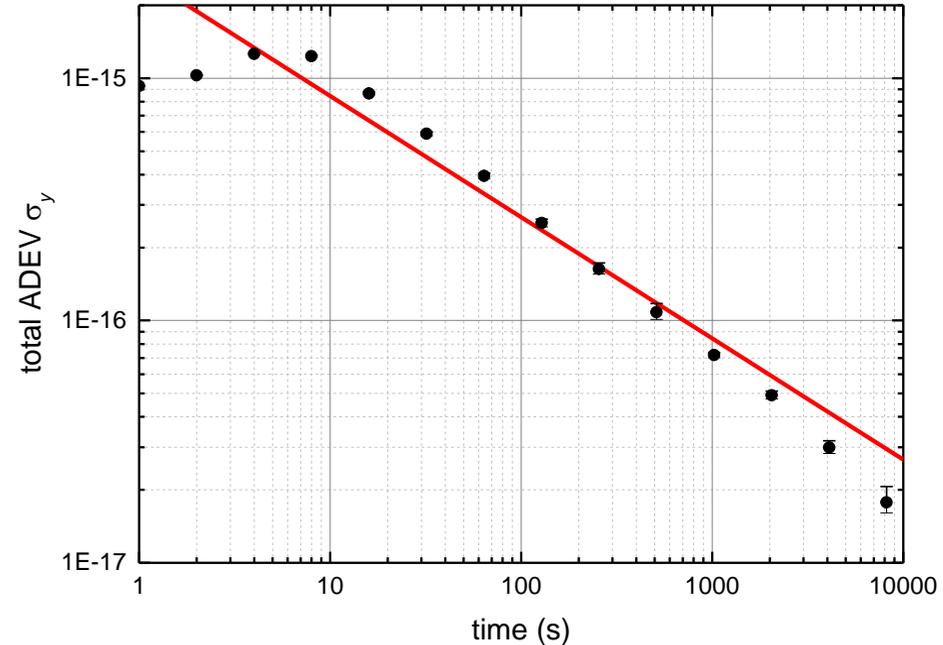
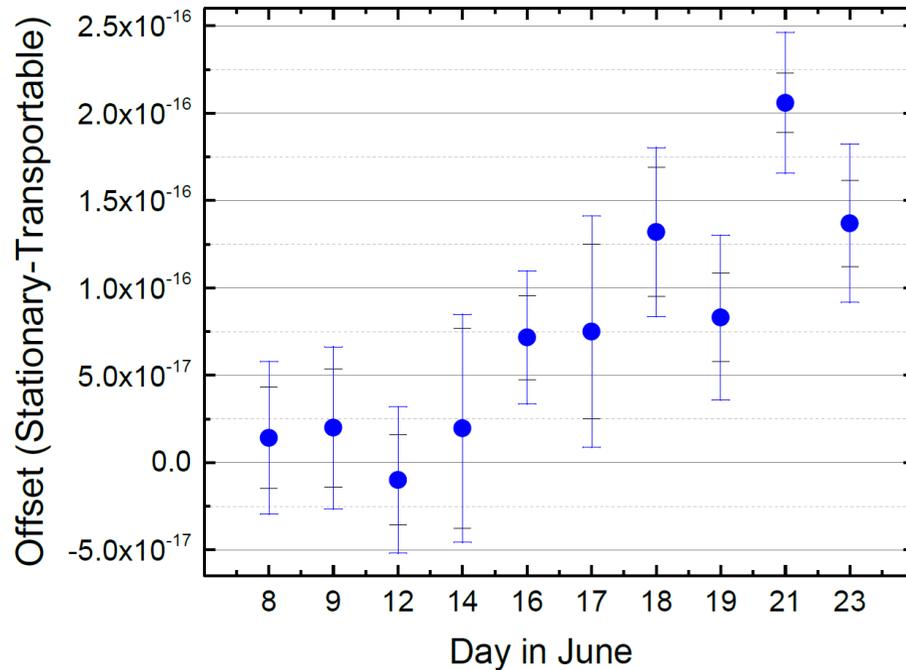
Second campaign: Paris – Braunschweig 2017



Combined uncertainty $\approx 3 \times 10^{-17}$ or 30 cm in 3 hours.

Gravity potential correction from geodesy: $-247.2(4) \times 10^{-17}$

Second campaign: Paris – Braunschweig 2017

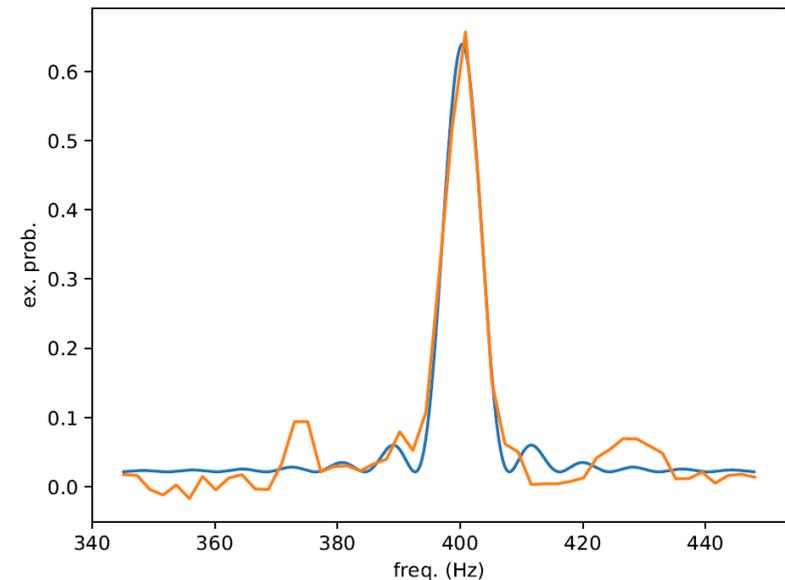


Combined uncertainty $\approx 3 \times 10^{-17}$ or 30 cm in 3 hours.

Gravity potential correction from geodesy: $-247.2(4) \times 10^{-17}$

unfortunately: ‘anomaly’ in the second half of the campaign

- Need for further practise and improvements
- 5 days after arrival: atoms in the lattice
- second week: problems with spectroscopy
perturbation of the clock laser



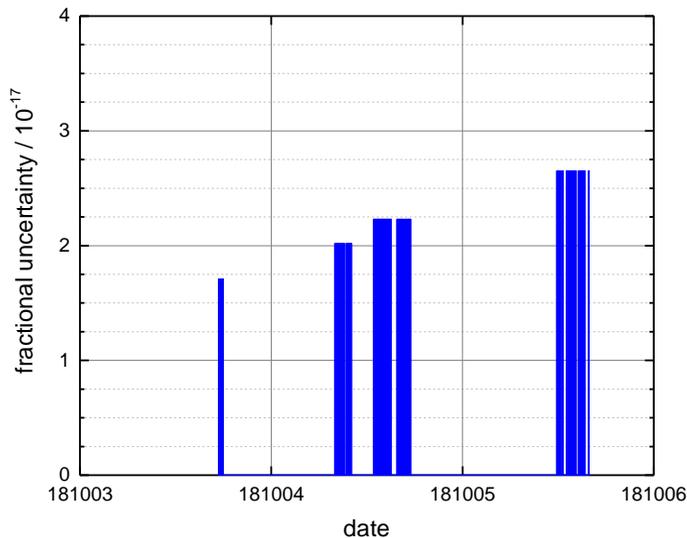
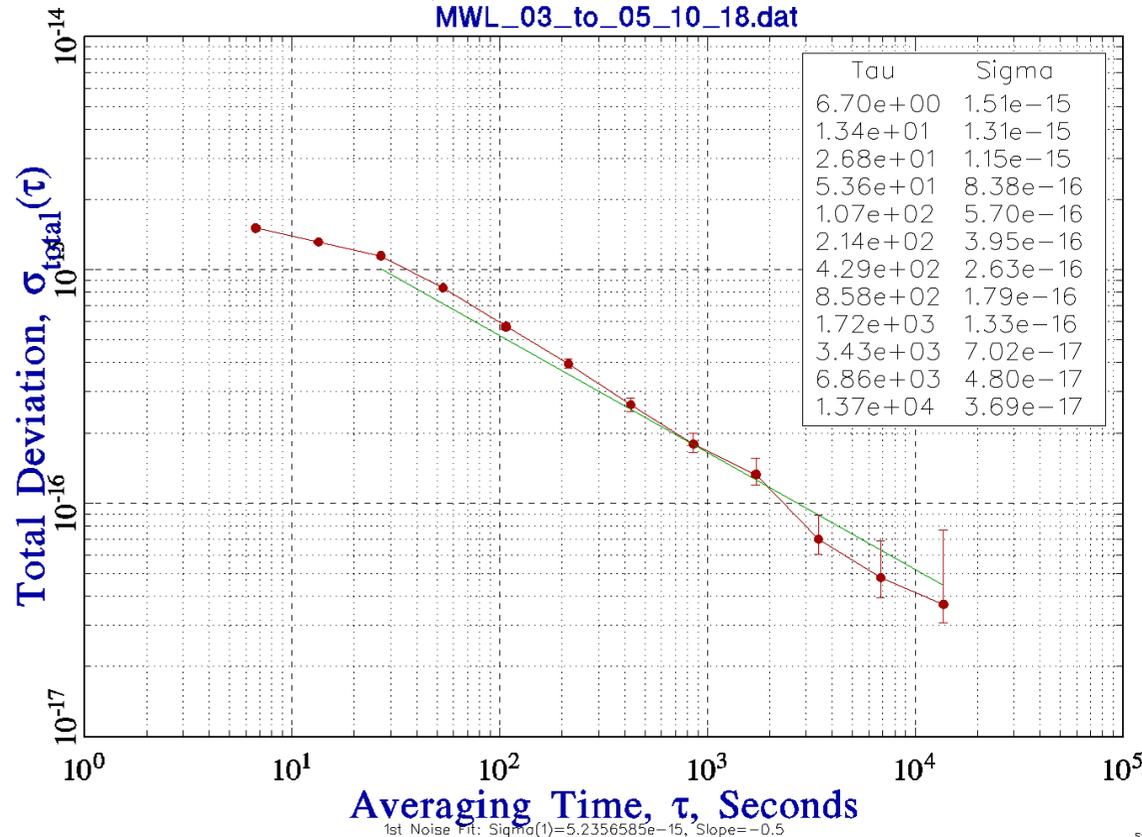
Third campaign: Munich – Braunschweig 2018

- Third week: Clock laser sidebands removed (fibre between clock laser and cavity)

- instability:
 $5 \times 10^{-15} \tau^{-1/2}$

Date: 10/09/18 Time: 17:01:00 Data Points 1 thru 4875 of 4875 Tau=6.7000000e+00 File: MWL_03_to_05_10_18.dat

FREQUENCY STABILITY MWL_03_to_05_10_18.dat

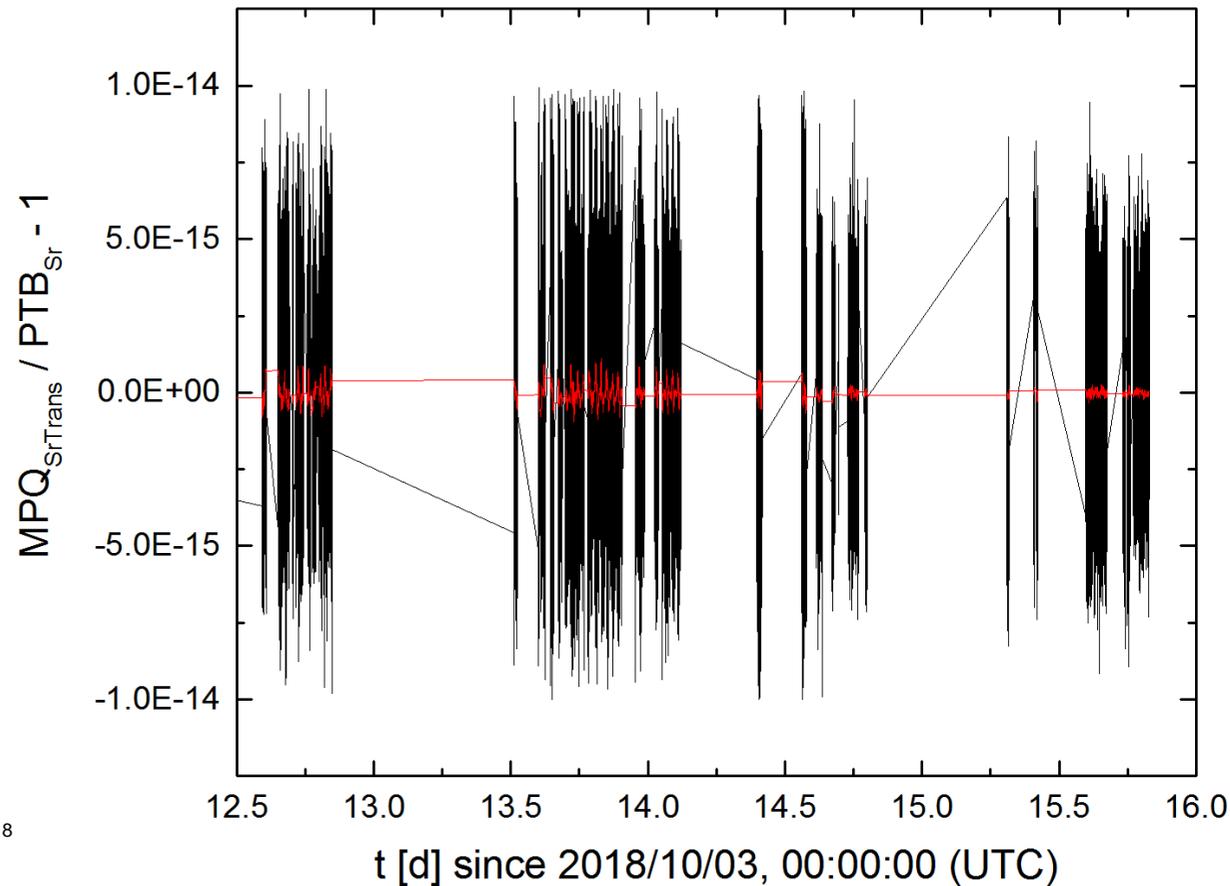
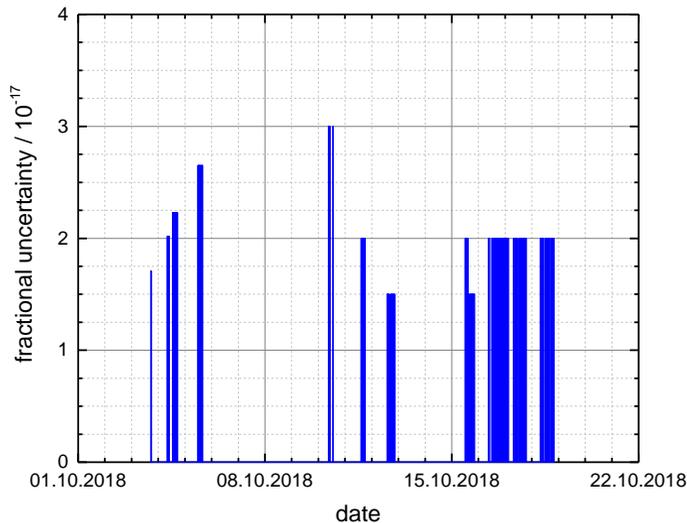


1st Noise Fit: Sigma(1)=5.2356585e-15, Slope=-0.5

Stable32

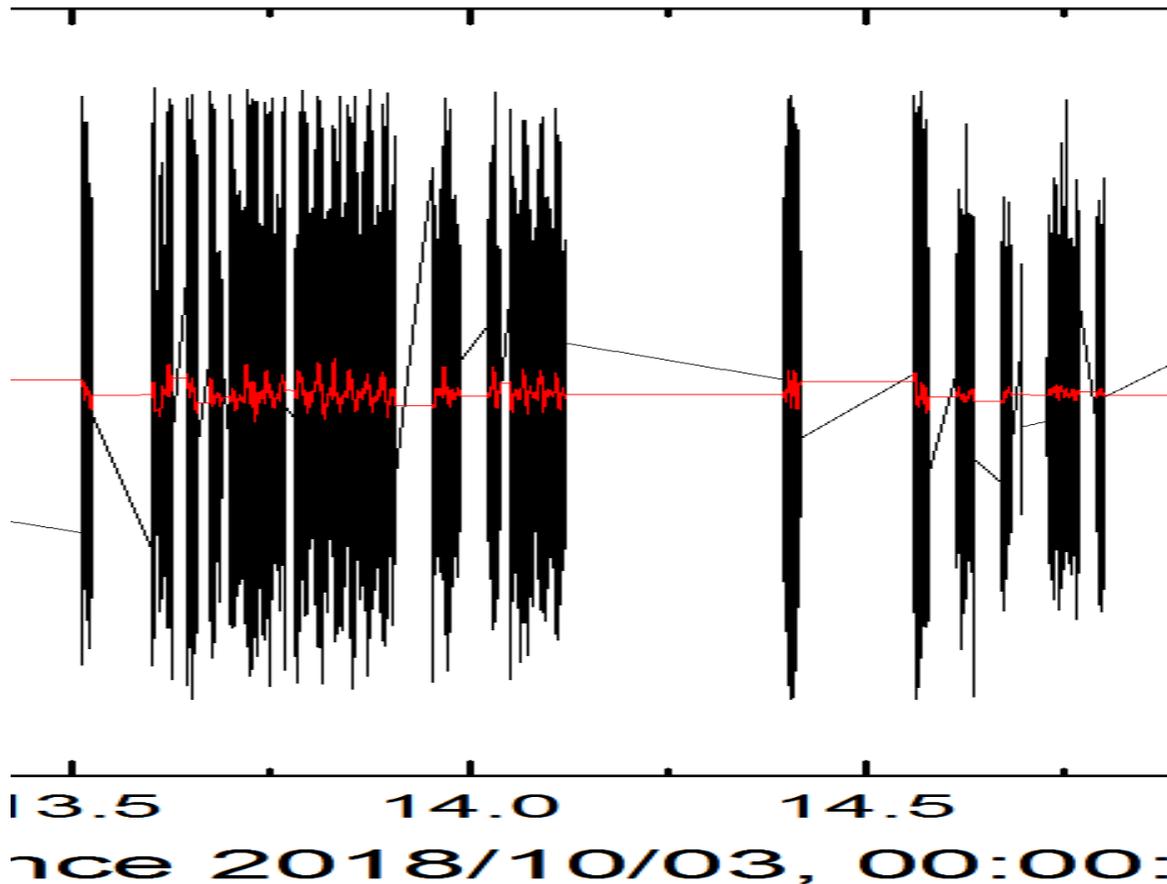
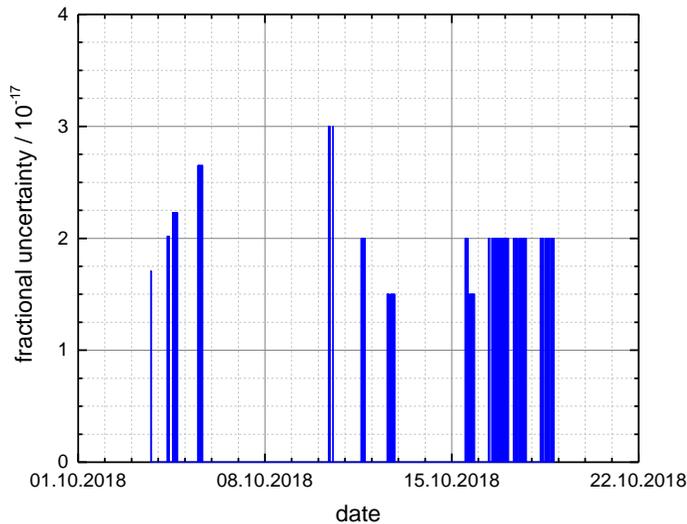
Third campaign: Munich – Braunschweig 2018

- improved counting of clock laser by the comb (week 4)
- more data ... (week 5)



Third campaign: Munich – Braunschweig 2018

- improved counting of clock laser by the comb (week 4)
- more data ... (week 5)



In conclusion:

- ▶ There is always something new ...
- ▶ Reliability is still an issue
procedures for 'quality management' have to be improved
hardware improvements are ongoing
- ▶ I still think that we can do this, even in space!
- ▶ 10^{-17} or 10 cm now, 10^{-18} or 1 cm in a few years

Many thanks to:

Strontium:

J. Grotti

S. Koller

S. Herbers

S. Vogt

S. Dörscher

A. Al-Masoudi

R. Schwarz

Cavities & Combs:

S. Häfner

E. Benkler

D. Matei

T. Legero

U. Sterr

Yb⁺ group at PTB

Fibre link group at PTB

Teams in NMIs

Italy (INRIM)

France (SYRTE)

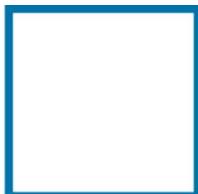
Teddington (NPL)

MPQ team at Munich

SOC team

& people who
want to join us





**Physikalisch-Technische Bundesanstalt
Braunschweig and Berlin**

Bundesallee 100

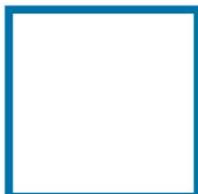
38116 Braunschweig

Christian Lisdat

Telefon: 0531 592-4320

E-Mail: christian.lisdat@ptb.de

www.ptb.de



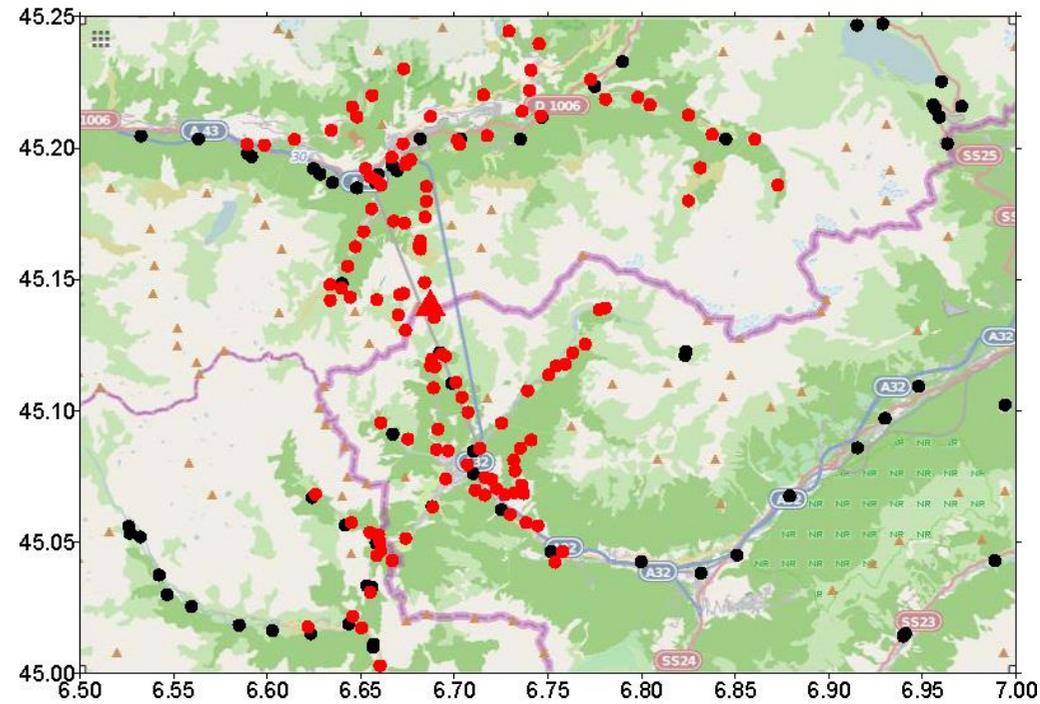
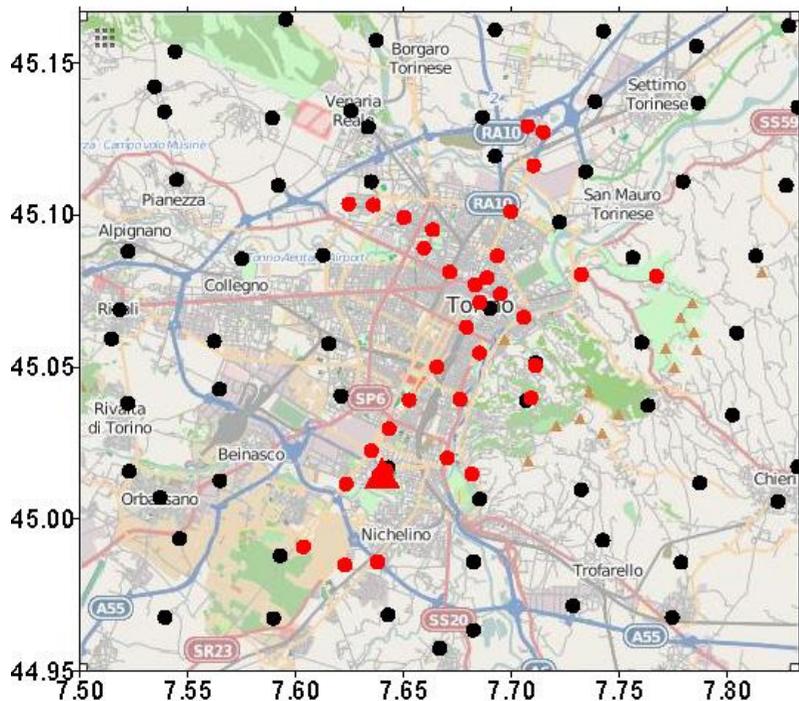
Stand: 02/17



- ▶ Approaching design uncertainty of 1×10^{-17}
- ▶ Reliability is still a problem
 - but it is obvious that you can do better
- ▶ Balance design/construction effort with salary of PhD student
- ▶ Next generation: lower uncertainty (1×10^{-18} ?)
more 'user friendly'
as heavy and power hungry



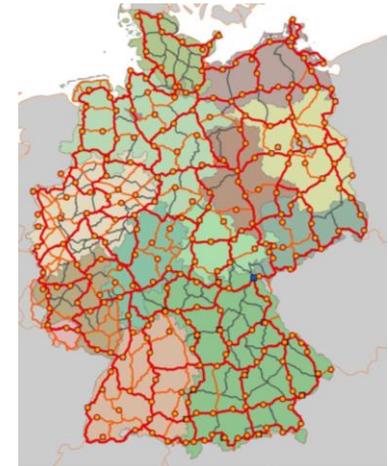
Gravity measurements Modane/Torino



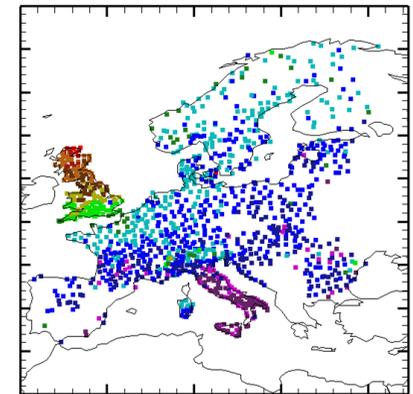
© OpenStreetmap

Erdbeobachtung:

- Nivellement
 - kleinschrittig (60 m)
 - Fehlerakkumulation
 - Schwerefeldmessung (Geoid) & GNSS
 - Satellitendaten
 - niedrige Ortsauflösung
- 3 cm Höhenauflösung über 500 km
 $\cong 3 \times 10^{-18}$ Uhrengenauigkeit



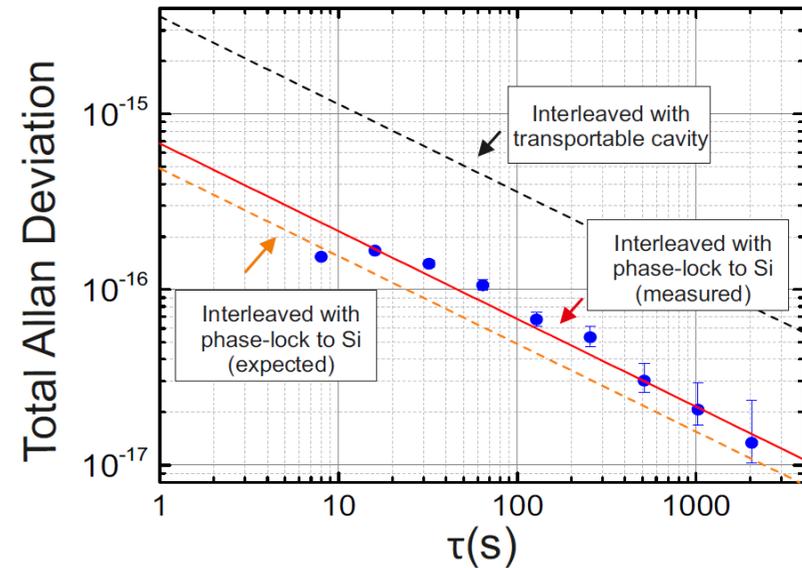
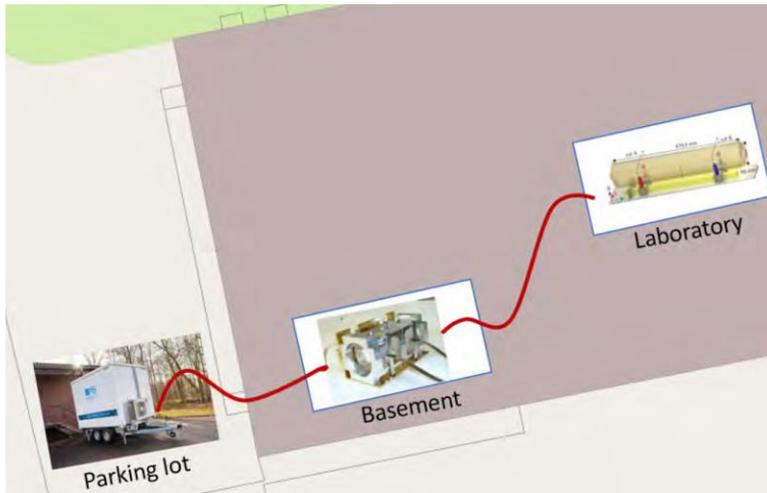
source: BKG/IfE



Levelled heights – GNSS/geoid heights (m)

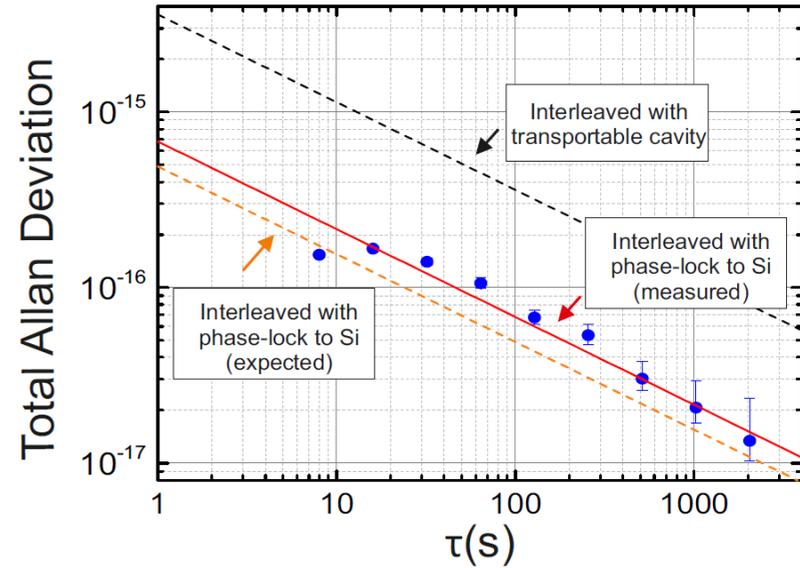
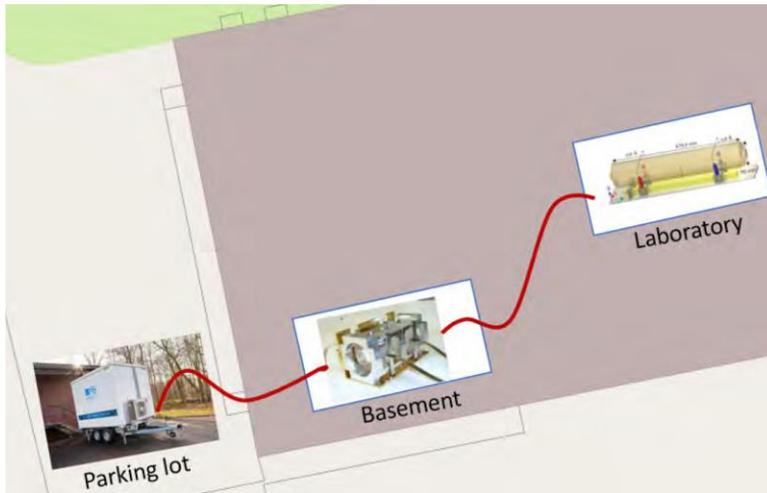
Gruber et al., ESA report GO-HSU-PL-0021
Height System Unification with GOCE

Further testing at PTB – know your clock



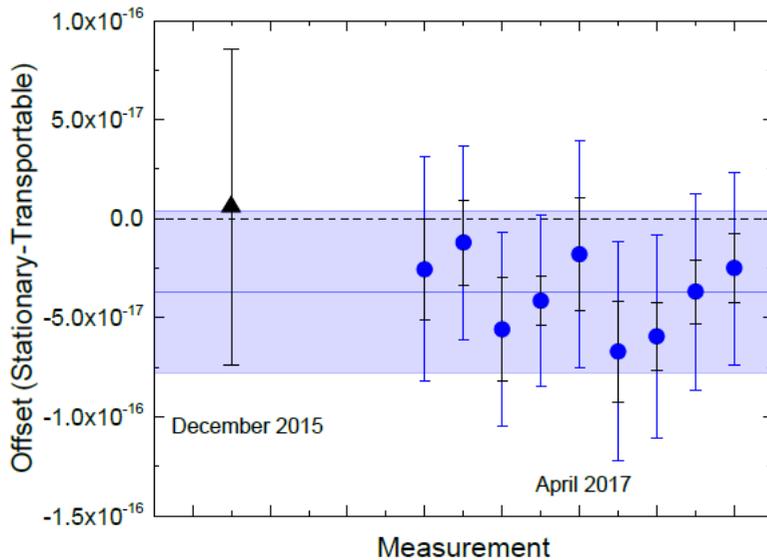
faster averaging using laboratory lasers

Further testing at PTB – know your clock

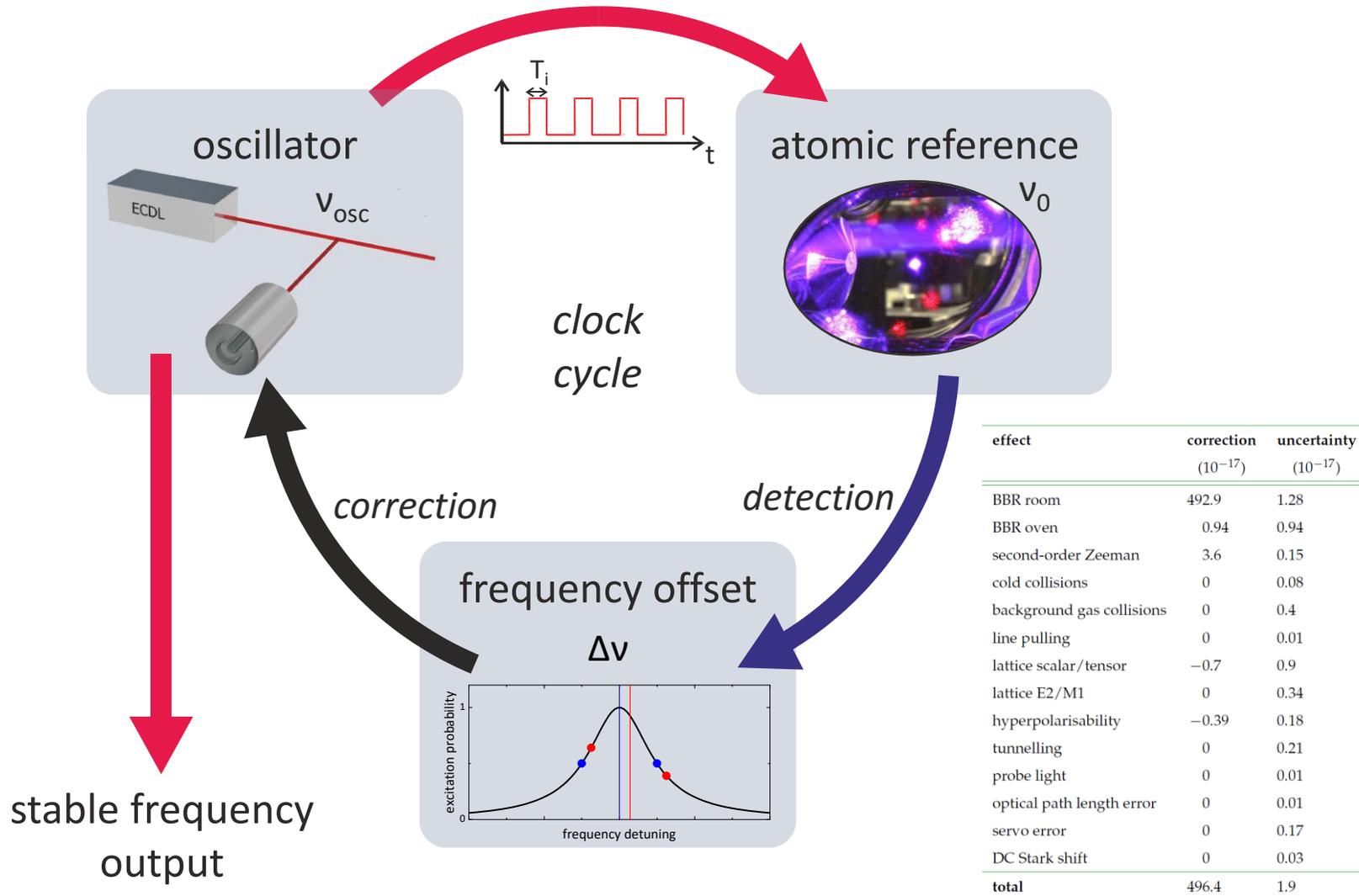


faster averaging using laboratory lasers

$$\nu_{\text{stat}}/\nu_{\text{trans}} - 1 = -37(41) \times 10^{-18}$$



Principle of operation



Control of collisional effects in 1D optical lattice

Increased MOT beams power (last step of the 2nd - stage MOT)



Higer temperature



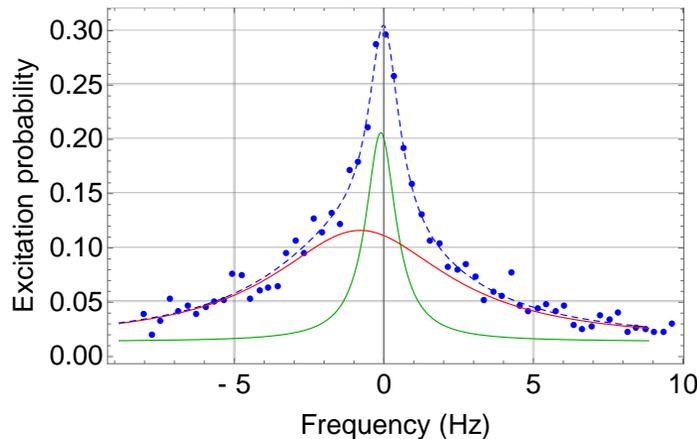
Lower density

The lattice power is then ramped down before interrogation in order to reduce the atoms' temperature.

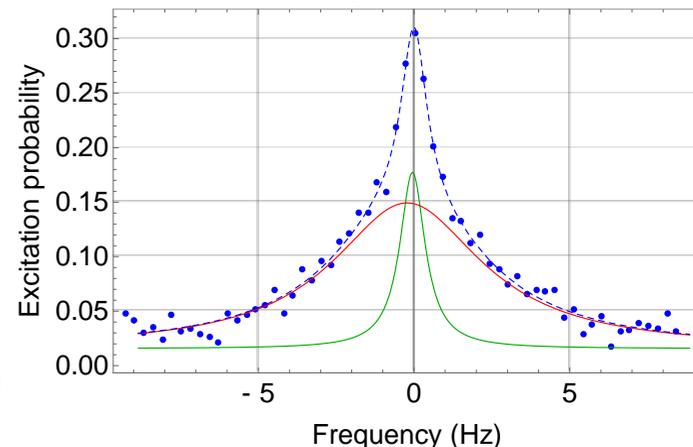
We fit the lineshape with the sum of two lorentzian curves:

- 1) **Green**: Fourier limited line → SINGLE OCCUPANCY
- 2) **Red**: broader line, red detuned → MULTIPLY OCCUPANCY

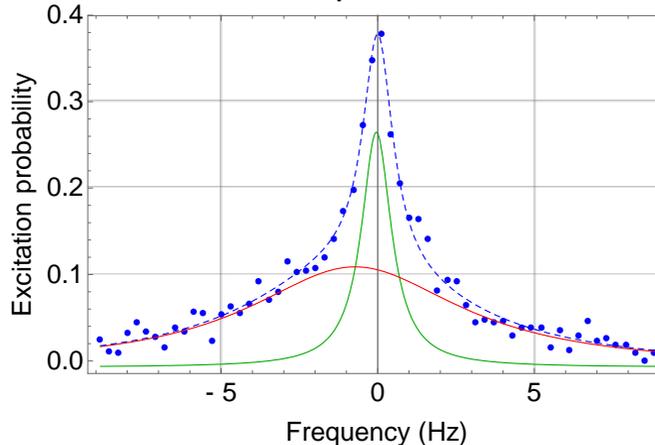
MOT beams power ~15 a.u.



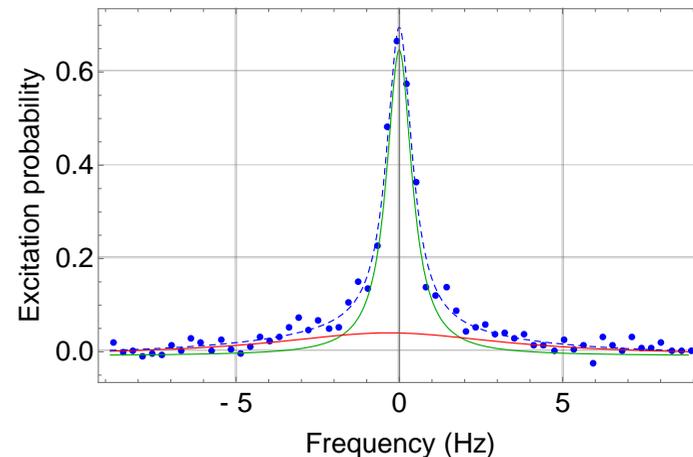
MOT beams power ~40 a.u.



MOT beams power ~90 a.u.



MOT beams power ~120 a.u.



Line pulling at 10^{-17} level, under investigation.

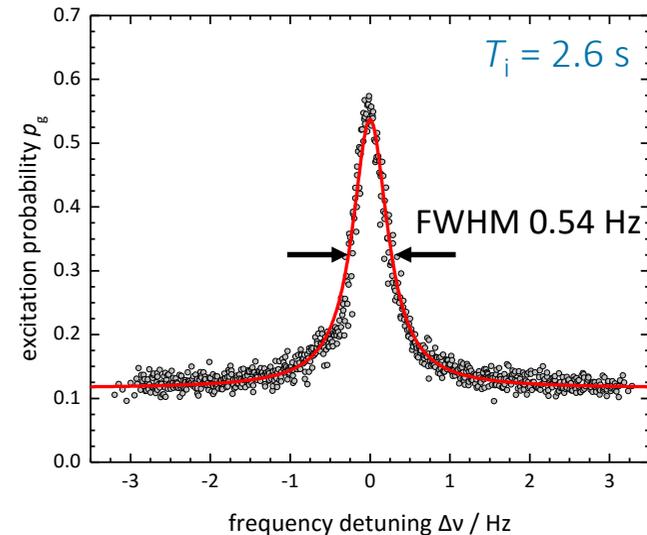
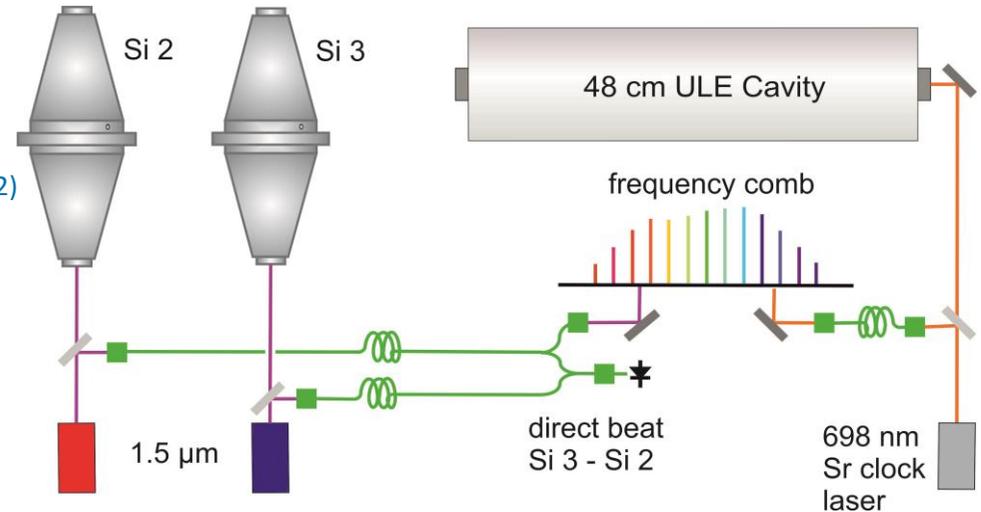
- ▶ Transfer stability from silicon cavities by stabilising beat.

H. R. Telle et al., Appl. Phys. B **74**, 1 (2002)

- ▶ Improved laser coherence

- Uninterrupted operation for 12 hours with 2.6 s interrogation
- Regular operation with ~ 1 s interrogation

- ▶ Improved stability



- ▶ Instability with ULE[®] resonator only:

$$1.6 \times 10^{-16} \tau^{1/2}$$

A. Al-Masoudi et al., Phys. Rev. A **92**, 063814 (2015)

- ▶ Instability with silicon resonators:

$$5 \times 10^{-17} \tau^{1/2}$$

- ▶ Instability with dead time-free interrogation:

$$6 \times 10^{-17} \tau^{1/2} \text{ @ NIST}$$

M. Schioppo et al., Nature Photonics **11**, 48 (2017)

