ТЛ

# Full-scale Monte-Carlo Data Simulation and Evaluation in the ELT Data Center

Christoph Bamann, Stefan Marz, Anja Schlicht, Rebekka Abel

Technical University of Munich

Chair of Satellite Geodesy

Mircufurm des TVM

ACES Workshop, Munich, 22-23 October, 2018



## ELT principle

Optical link, pulsed





## Short introduction to ELT (continued)

Principle of ELT (optical link, pulsed)

- One way:

$$tof_{1W} = R_{COM} + \tau_{troposphere} + \tau_{Sagnac} + \tau_{Shapiro} + \tau_{attitudeDetector}$$

- Two way:

$$tof_{2W} = 2 * \left( R_{CoM} + \tau_{troposphere} + \tau_{Shapiro} + \tau_{attitudeReflector} \right) + \tau_{Reflector}$$

with  $R_{COM}$ : Distance between spacecraft CoM and station reference

- Time transfer:

$$\tau = \frac{t_{return} + t_{start}}{2} - t_{detector} + \tau_{corr} = \frac{t_{of_{2W}}}{2} + t_{start} - t_{detector} + \tau_{corr}$$



## Capabilities of the simulation environment

#### **Geometric components**

- Earth orientation (IERS 2010 Conventions)
- ISS attitude simulation
   (3 axes, constant offsets and oscillations)
- Detector and reflector position
- Intra-reflector delay (function of incidence angle)
- Visibility constraints (minimum elevation

#### Signal delays

- Troposphere (including cloud cover)
- Sagnac effect (processing in ITRF)
- Shapiro delay

#### **Relativistic effects on clocks**

- Drift of clocks w.r.t. to UTC
  - ... due to special relativity (relative velocity)
  - ... due to different gravitational potential

#### **Stochastic components**

- Background noise
- Laser Jitter
- Pulse width
- Noise of ground- and space-based clocks

# Capabilities of the simulation environment (continued)

Station	_ Orbit	- Attitude (right-handed, z in padir direction)
D         Name           8834 (wetl)         Wettzell	Precise-filename select (.mat) ISS_2017-04-03_8834_Ts1_pass15_full	X [m]     Y [m]     Z [m]       Reflector offset     10.778     9     7.021   Select reflectors ELT reflector
X [m] 4075576.613	Orbit errors in predictions (CPF) cr0 cr1 cr2 Radial [m]	X [m]         Y [m]         Z [m]         JEM Hemi A           Detector offset         10.978         9         7.021         JEM Hemi B         Image: Comparison of the mini B
Y (m) 931785.727 Z (m) 4801583.739	Call         Call <th< td=""><td>Satelite         Angle         ∆r const.         ∆r∆r         ∆r∆r         C2V2 S3 Forward         ✓           rotation         [deg]         [deg]         [deg]         period [s]         type         C2V2 S3 Forward         ✓           Roll (x)         0.7         0.5         0         0         ✓ sin</td></th<>	Satelite         Angle         ∆r const.         ∆r∆r         ∆r∆r         C2V2 S3 Forward         ✓           rotation         [deg]         [deg]         [deg]         period [s]         type         C2V2 S3 Forward         ✓           Roll (x)         0.7         0.5         0         0         ✓ sin
Station Parameters Station Range Bias [m]	Cross-track [m]         0         0         0           Orbit time bias [ns]         0         0         0         0	Pitch (y)0.40.5 0 0 ☑ sin
Temperature [K] 288.4	Clock Parameters	
Pressure [hPa] 947.4 Rel. Humidity [%] 86	Station clock offset to UTC [s] Uter [ns^2] Uter [ns^2] Uter [ns^2] Uter [ns^2]	Detector parameters           Jitter [ns'2]         Gate width [ns]         Activation time [ns]           0.02         0.0014         100         0
Calibration Offset [ns] 9999 Two-way Detection Probability 0.1	ACES clock offset to UTC [s]	One-way detection Background Minimum elevation probability noise rate [1/s] for visibility [deg]
Jitter [ns^2] 0.0014	0 Zenith gap ☑ Minimum elevation zenith gap [deg]	n 80
Background noise rate [1/s] Background	ACES: SHM + PHARAO V	Load defaults Save inputs as new defaults
	Write simu	ulation files
Cloud cover [%] Cloud width [s]	Optional folder end Release(s) (0 to 9 for storing the files	99) 0 Evaluation Start



## ELT data processing

ELT reflector identification and coarse 2-way filtering

Final 2-way filtering and orbit (attitude) improvement

1-way filtering based on orbit (attitude) corrections

	Preliminary noise reduction	
	Build histogram	
$\Rightarrow$	Exponential noise reduction	
	Signal peak localization	
	Iterative sigma screening	

QUICKIOOK	
Evaluation folder	
8834-20170401-170537-20170401-	171506 select folder
Peeidual pre_fittering threshold [pe]	
Residual pre-filtering threshold [ns] 3.0	Write PDF file
Residual pre-filtering threshold [ns] 3.0 Filtering sigma-factor [-]	Vrite PDF file



#### Simulation of the ACES clock







## Signal delay in the ELT reflector

#### **Description of ELT CCR Configuration**

To extract ranges from SLR measurements and to consider geometrical corrections the configuration of the CCR has to be known. This implies <u>knowledge of the location of the CCR reference point</u> with respect to the ACES coordinate system, <u>the orientation of the CCR and the CCR correction table</u>.

#### ELT CCR reference point and orientation

ASCII file containing list of start time and ELT CCR reference point and orientation information. The ELT CCR <u>detector reference point is defined in relation to the ACES reference point</u> and the <u>orientation of the CCR is</u> <u>defined by quaternions</u>. The x-axis of CCR is in azimuth = 0, y-axis in azimuth = 90° and z-axis in boresight.

#### **CCR** Correction

<u>Azimuth and elevation-dependent range corrections relative to the mechanical reference point</u> of CCR have to be taken into account. These corrections <u>imply the optical and mechanical parameters of the reflector array</u>. The corrections are <u>provided in the nearest-prism approximation</u>.



#### Elevation cut-off, zenith gap, cloud coverage (continued)



- Clouds of equal "duration" and regular distribution
- <u>Parameters</u>: Total cloud coverage  $P_{clouds}$  [%/pass] and cloud frequency  $f_{clouds}$  [#/pass]

#### Elevation cut-off, zenith gap, cloud coverage (continued)



Simulated 2-way range residuals for different cloud coverage parameters

Christoph Bamann | ACES Workshop 2018 | October 22-23



#### Cloud coverage and periodic attitude error





## Multiple reflectors on the ISS



Name	X [m]	Y [m]	Z [m]
JEM LRR Hemi A	10.878	-5.448	7.021
JEM LRR Hemi B	10.876	-6.092	7.017
IDA 1 Hemi B (a)	15.789	0.891	6.239
C2V2 S3 Forward Antenna Boom Hemi (c)	1.524	22.887	-1.417
C2V2 P3 Nadir Antenna Boom Hemi (c)	-2.621	-22.887	-0.978



## Multiple reflectors on the ISS (continued)

For reflectors  $i = 0 \dots N$  with distances  $d_i \le d_{i+1}$  with respect to the observer the probability of detecting a signal is:

$$p_{i,eff} = p_i \prod_{j=0}^{i-1} (1-p_j)$$

The single-reflector probabilities  $p_i$  may account for differences in the effective cross section among the reflectors.

With a constant background noise rate, the noise statistics follows an exponential distribution in single photon mode:

$$p_{signal.i} = p_{eff,i} e^{-n_{noise}\Delta t_i}$$





### Multiple reflectors on the ISS (continued)





#### New noise reduction algorithm

$$P_{s>thr} = \sum_{m=N_t}^{\infty} p_{binomial,n+s}(m) \qquad P_{n+s}(T_x) = 1 - \exp\left[-(n_n T_x + N_s)\right]$$

$$p_{binomial,n+s}(m) = \frac{y!}{m!(y-m)!} P_{n+s}(T_x)^m (1 - P_{n+s}(T_x))^{y-m}$$

$$P_{n>thr} = \sum_{m=N_t}^{\infty} p_{binomial,n}(m) \qquad P_n(T_x) = 1 - \exp\left[-n_n T_x\right]$$

$$p_{binomial,n}(m) = \frac{y!}{m!(y-m)!} P_n(T_x)^m (1 - P_n(T_x))^{y-m}$$





#### New noise reduction algorithm





## Monte-Carlo simulations

- Data simulation and processing for identical parameters
- Passes

-

-

- Laser system characteristics
- Signal propagation characteristics
- ... (neglecting multiple reflectors on the ISS)
- Randomness introduced by the following sources
- Background noise
- Laser jitter
- Pulse width
- Clock noise

- Studies without systematic errors
  - Expected to converge to "true" clock offset
  - ... if filtering does not fail
  - ... and yields unbiased time transfer triplets
  - How does filtering perform statistically?
- Studies with systematic errors
  - Unknown <u>attitude</u> and orbit errors will be present (particularly in quick-look processing)
- Effects of <u>cloud coverage</u> and other constraints on performance



#### Monte-Carlo simulations (continued)

Background	Noise	Time transfer	
noise rate [1/s]	reduction	σ [ps]	
5.00E+05	no	1.6979	
5.00E+05	yes	0.4541	
5.00E+06	no	4.7743	
5.00E+06	yes	1.8626	



# Monte-Carlo simulations (continued)

Attitude error (only pitch)





## Monte-Carlo simulations (continued)

#### Attitude error (only roll)





## Outlook

- Further Monte-Carlo simulations will be performed (suggestions are welcome!)
- We identified the ISS attitude as a critical source of error (insights via ISS tracking campaign would be very helpful)
- Challenges due multiple reflectors (see next talk by Stefan Marz)